



The study of the ${}^7\text{Li}(\gamma, t){}^4\text{He}$ reaction with mono-energetic gamma-ray beams

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Young Researchers & Young Engineers Days - 2024

MOTIVATION OF THE EXPERIMENT

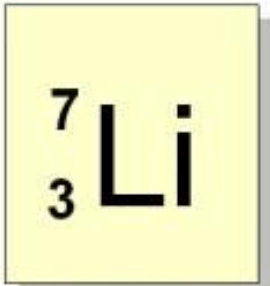
Big Bang Nucleosynthesis (BBN) → predicts the abundances of the light elements (produced in the early stages of the Universe)

“Cosmological Li problem”:

The measured ${}^7\text{Li}$ abundance is 3-4 times lower than expected



Observations on the low-metallicity stars



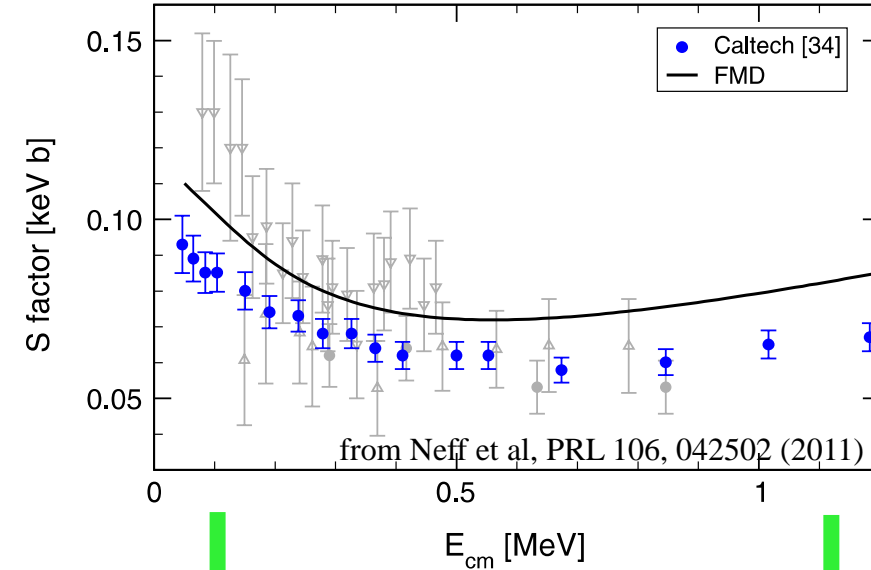
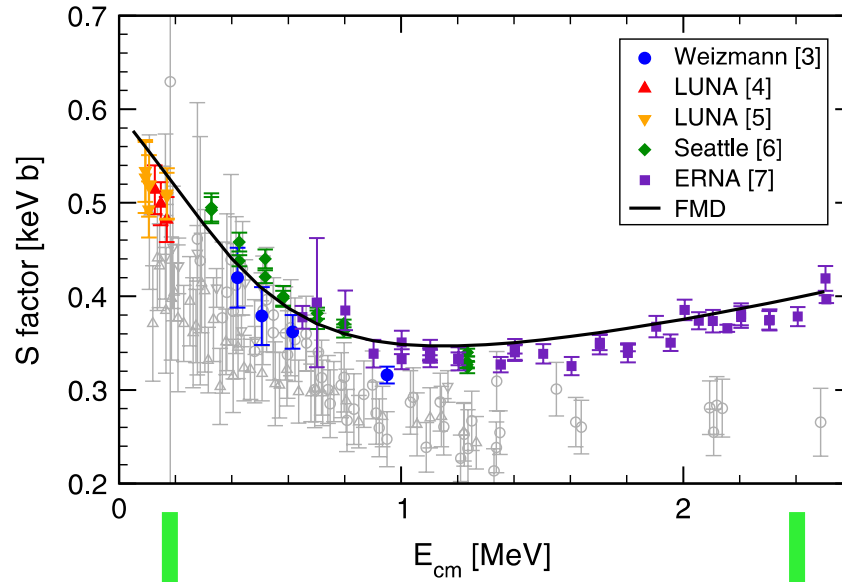
- The stellar measurements are leading to anomalous results
- OR**
- An error is present in the nuclear theoretical models



The error can be related to the WMAP baryonic density

Lower value → higher effect of the ${}^3\text{H}(\alpha, \gamma) {}^7\text{Li}$ reaction.

Two main reactions are responsible for the production of mass 7 elements:



FMD (fully microscopic fermionic molecular dynamics approach) doesn't fit the experimental data for ${}^3\text{H}(\alpha, \gamma) {}^7\text{Li}$

→ ${}^3\text{H}(\alpha, \gamma) {}^7\text{Li}$: only a few experiments were performed

→ Due to restricted use & health/safety, an experiment with a tritium target cannot be performed anymore

However, the ${}^3\text{H}(\alpha, \gamma) {}^7\text{Li}$ reaction can still be studied by its **INVERSE REACTION**

Reciprocity Theorem:

Inverse reaction



$$\frac{\sigma_{Bb \rightarrow Aa}}{\sigma_{Aa \rightarrow Bb}} = \frac{(2j_A + 1)(2j_a + 1)k_{Aa}^2(1 + \delta_{Bb})}{(2j_B + 1)(2j_b + 1)k_{Bb}^2(1 + \delta_{Aa})}$$



Direct reaction

*Photons: $2j_\gamma + 1 = 2$

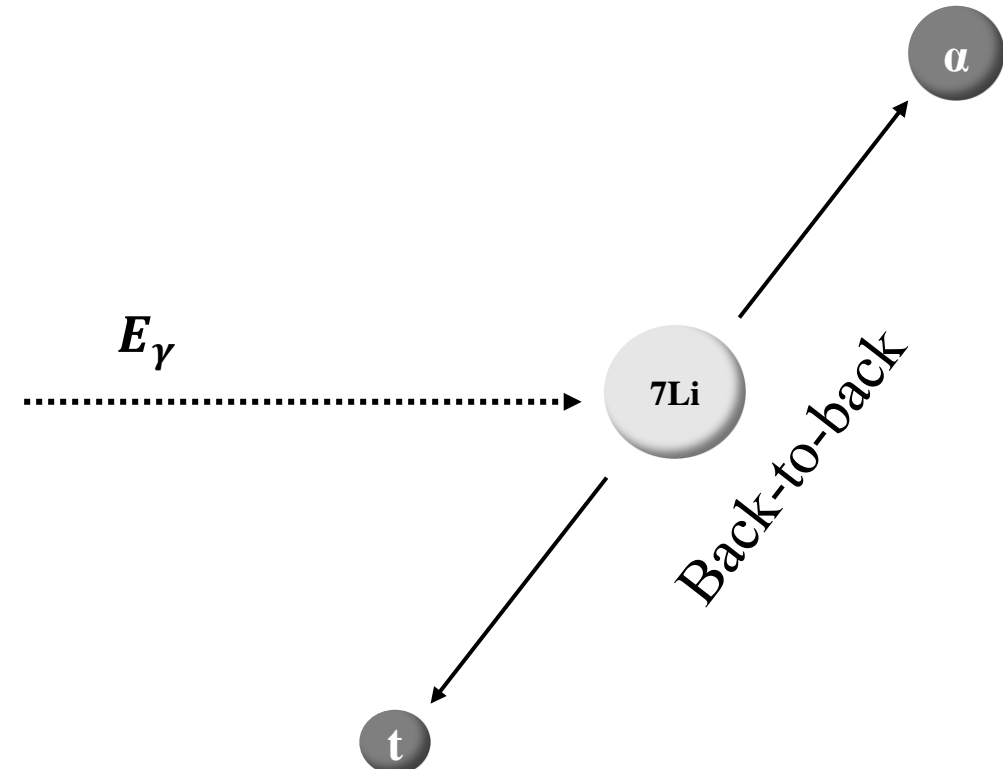
\Rightarrow

$$\sigma_{Bb \rightarrow Aa} = k \cdot \sigma_{Aa \rightarrow Bb}$$

${}^3\text{H}(\alpha, \gamma) {}^7\text{Li}$ \rightarrow Direct reaction

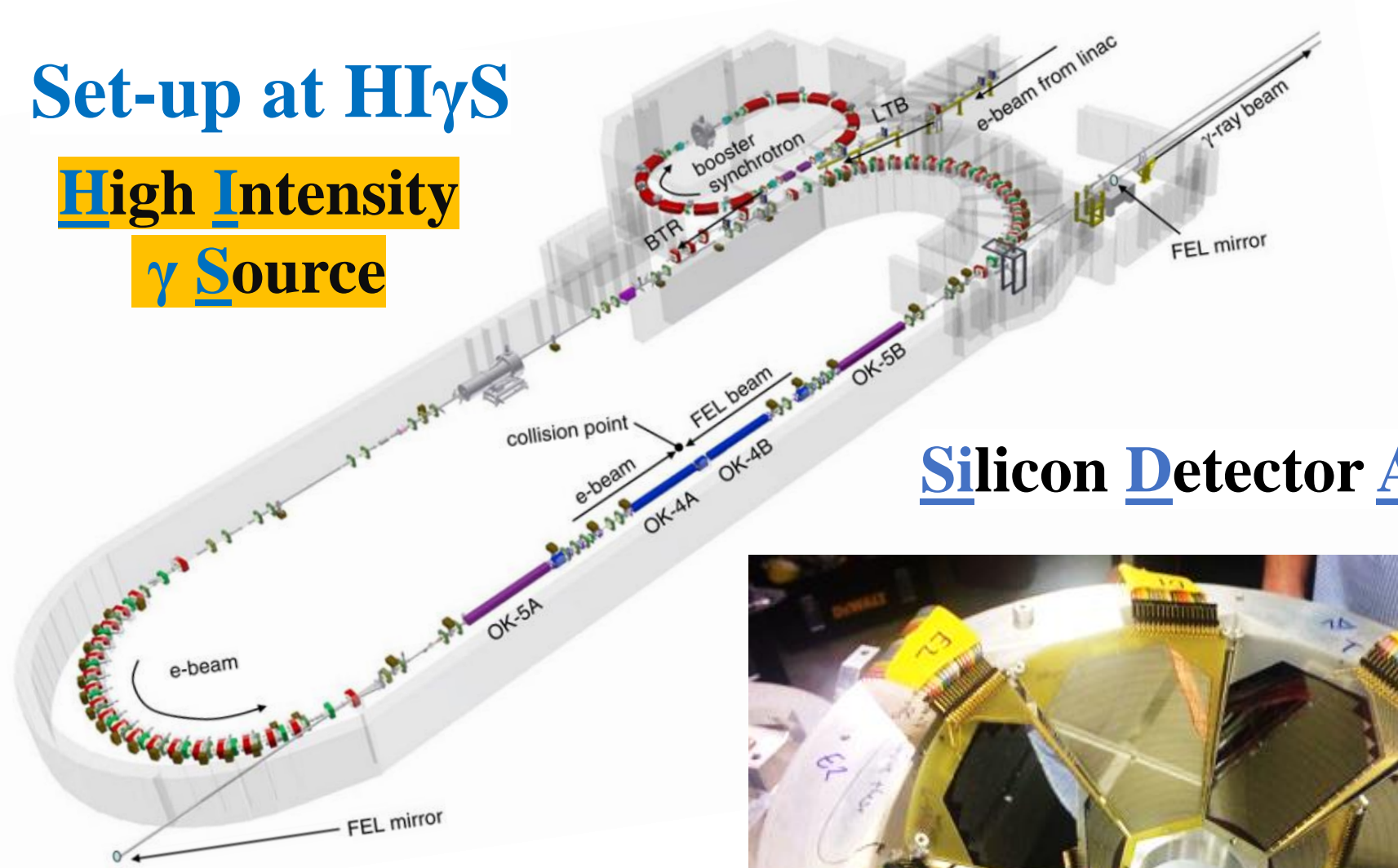
${}^7\text{Li}(\gamma, t) {}^4\text{He}$ \rightarrow Inverse reaction

\longleftrightarrow Lithium photodisintegration:
 $k \approx 60$

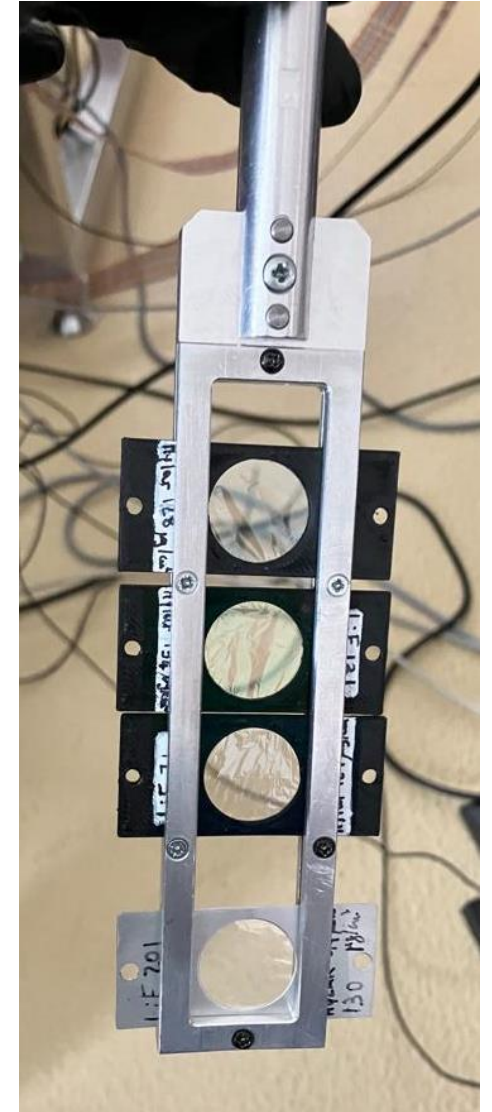
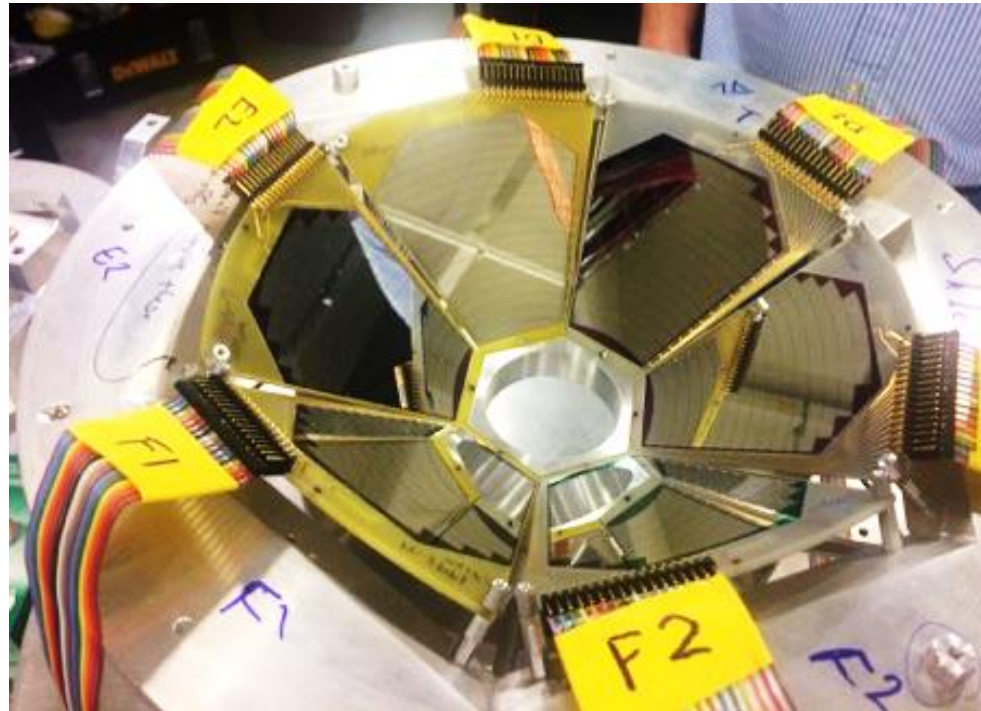


Set-up at HIγS

High Intensity γ Source



Silicon Detector Array



Energy:
1 MeV -100 MeV

LiF/Mylar

SIDAR



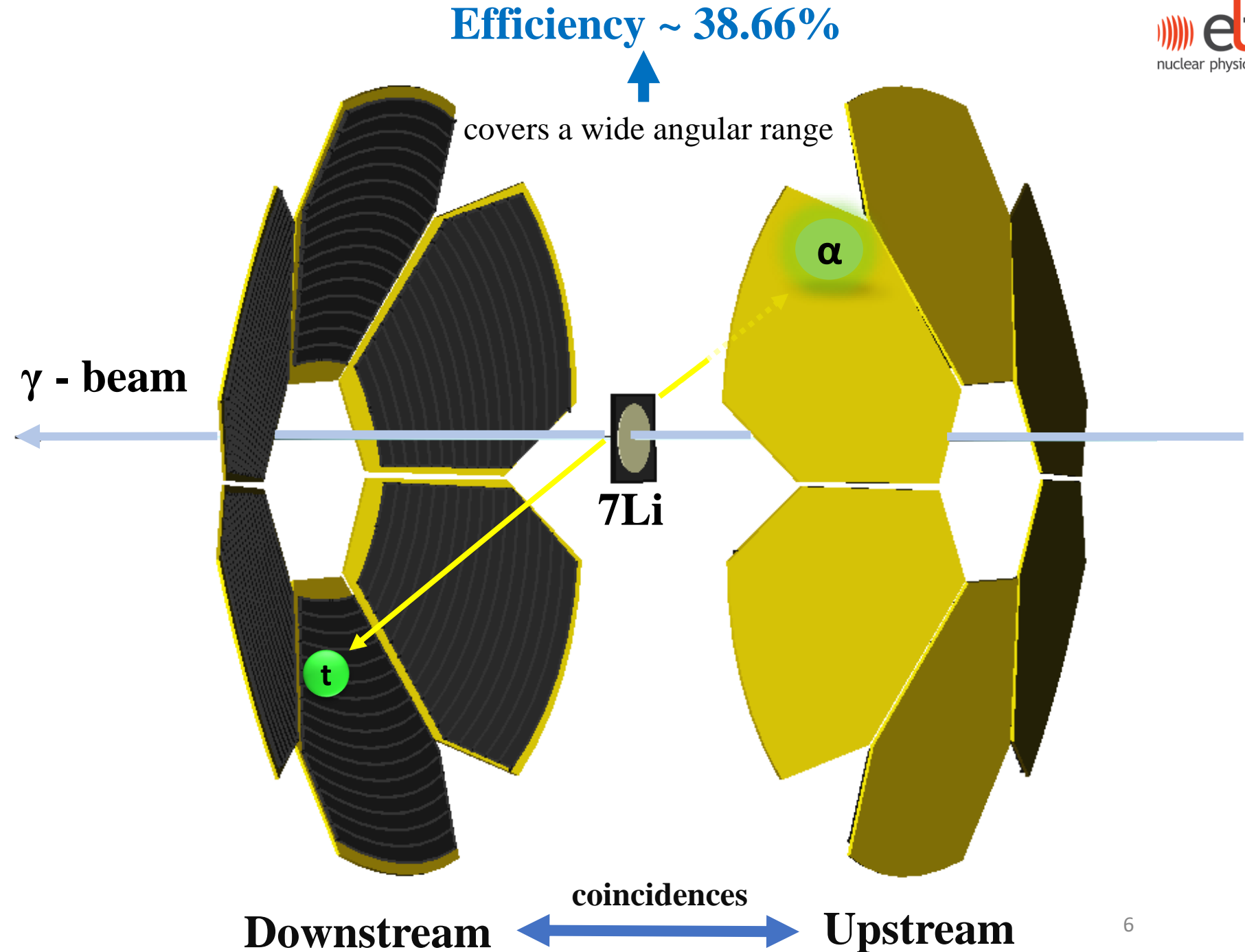
**Silicon
Detector
Array**



**2 x 6 YY1
(16 strips)**



192 channels

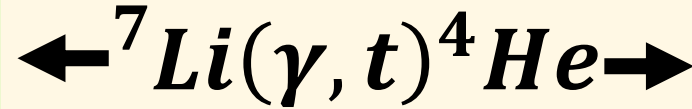


The old experiment:

2017

4.4 MeV - 10 MeV

energies above 6 MeV/thinner detectors



The new experiment:

6th April 2023-12th April 2023

3.7 MeV - 6 MeV

IMPROVEMENTS

1. ARRAY OF THINNER Si DETECTORS: 100&65μm

2. THINNER ENTRY FLANGE

**3. REDUCED ELECTRON DENSITY IN THE FLANGE
(Al instead of stainless steel)**

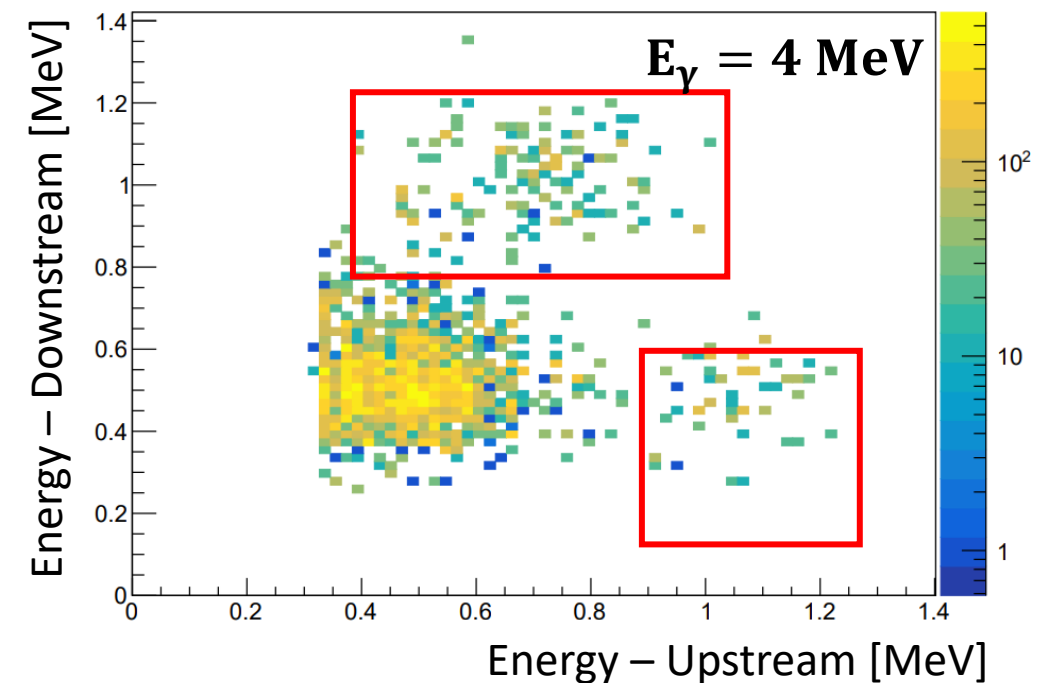
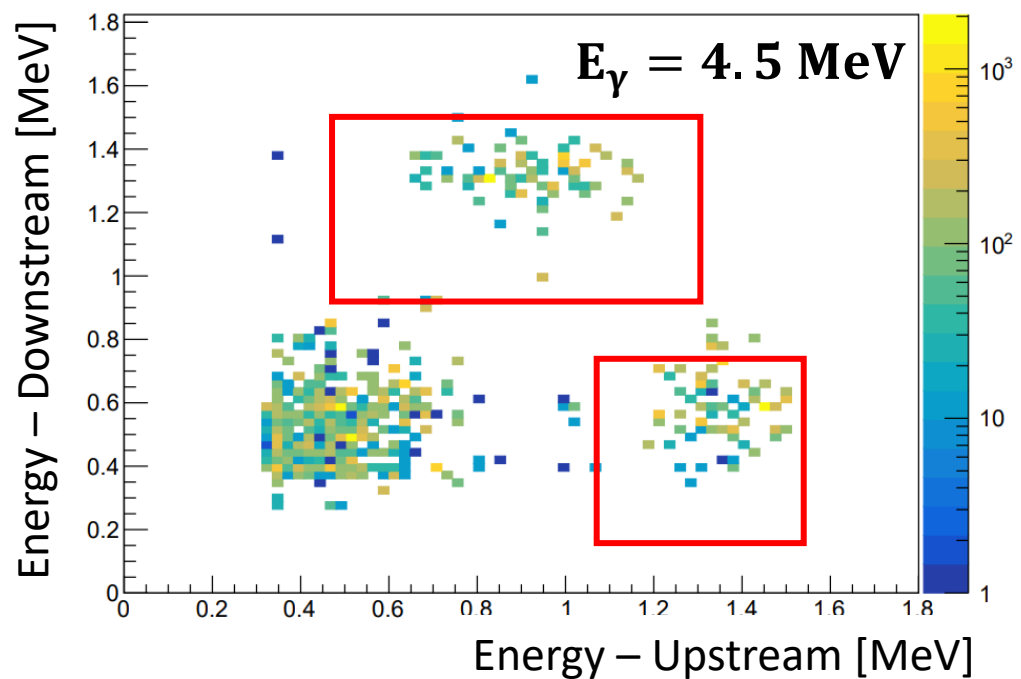
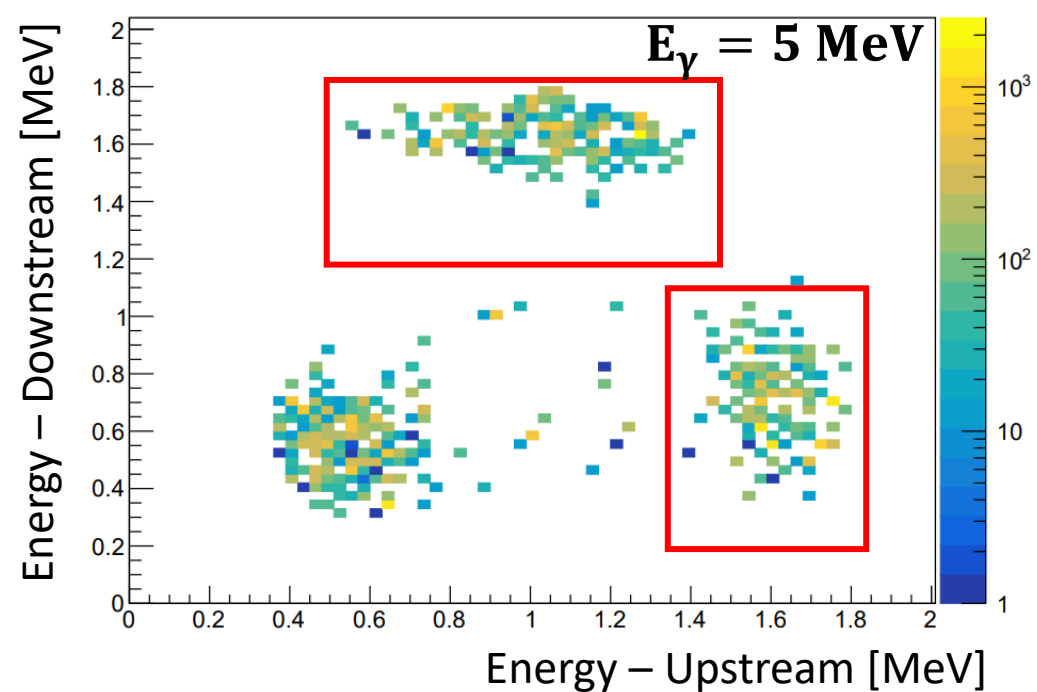
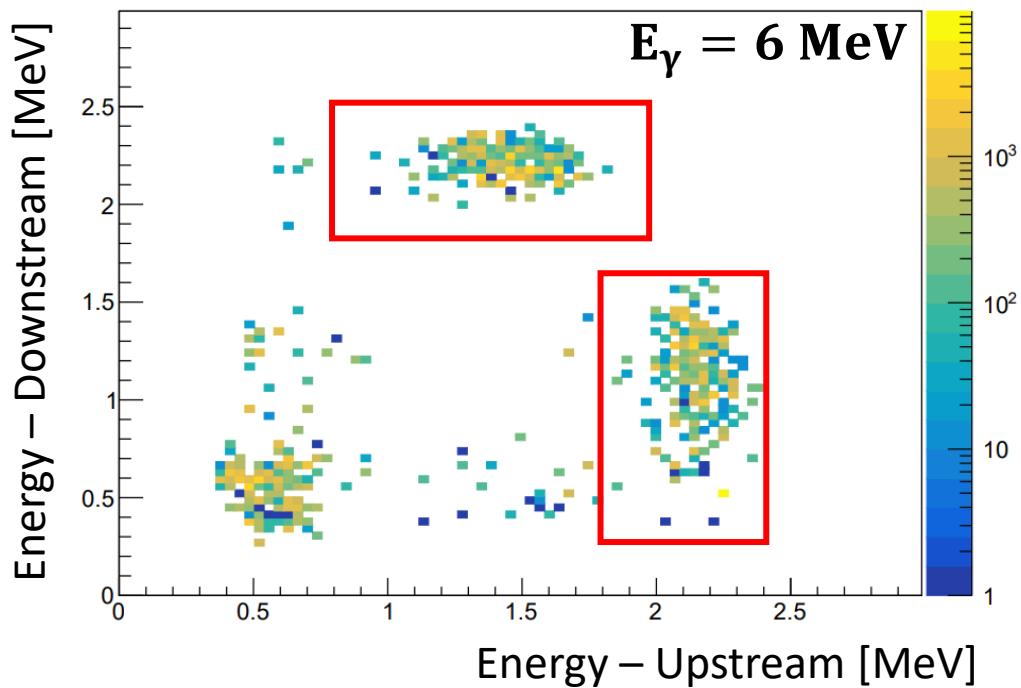
4. LONGER VACUUM PIPE (in front of the chamber)



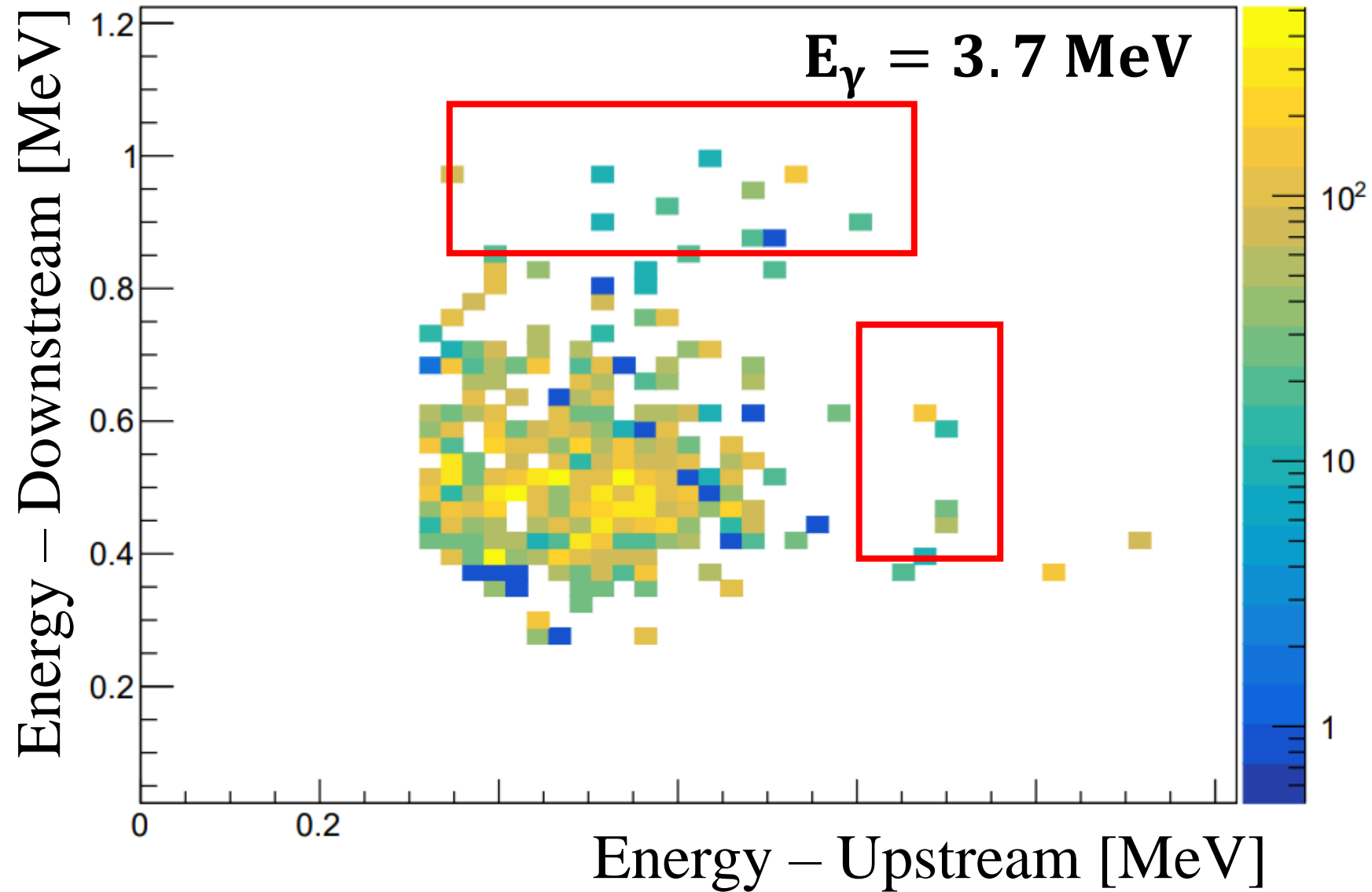
**The coincidences have been
clearly separated even for the
lowest energy**



**The background has been
highly reduced**



PRELIMINARY RESULTS



**Number of
coincidences**



R

$$\sigma = \frac{R}{N_{beam} \cdot N_{target} \cdot \varepsilon}$$



$$N_b = I \cdot t$$

where $I \sim 10^7 \text{ s}^{-1}$

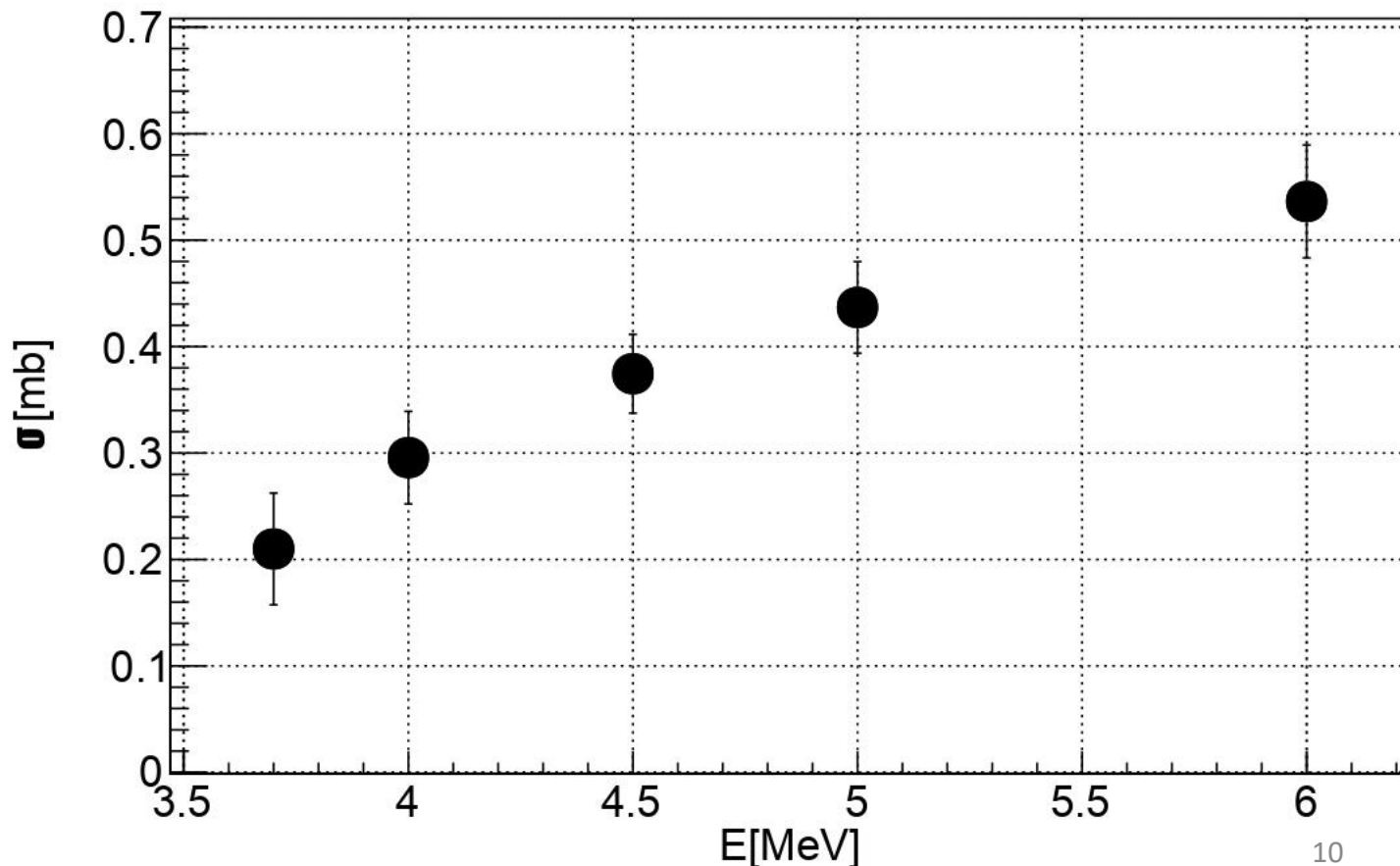


$$N_t \sim 10^{18}$$

Quantity	Uncertainty
Events no	~ 5-18 %
Intensity	~ 10%
Li-7 atoms no	~ 1%
Efficiency	~ 1%

The ground state cross-section of the
inverse ${}^7\text{Li}(\gamma, \alpha) {}^3\text{He}$ reaction

PRELIMINARY RESULTS

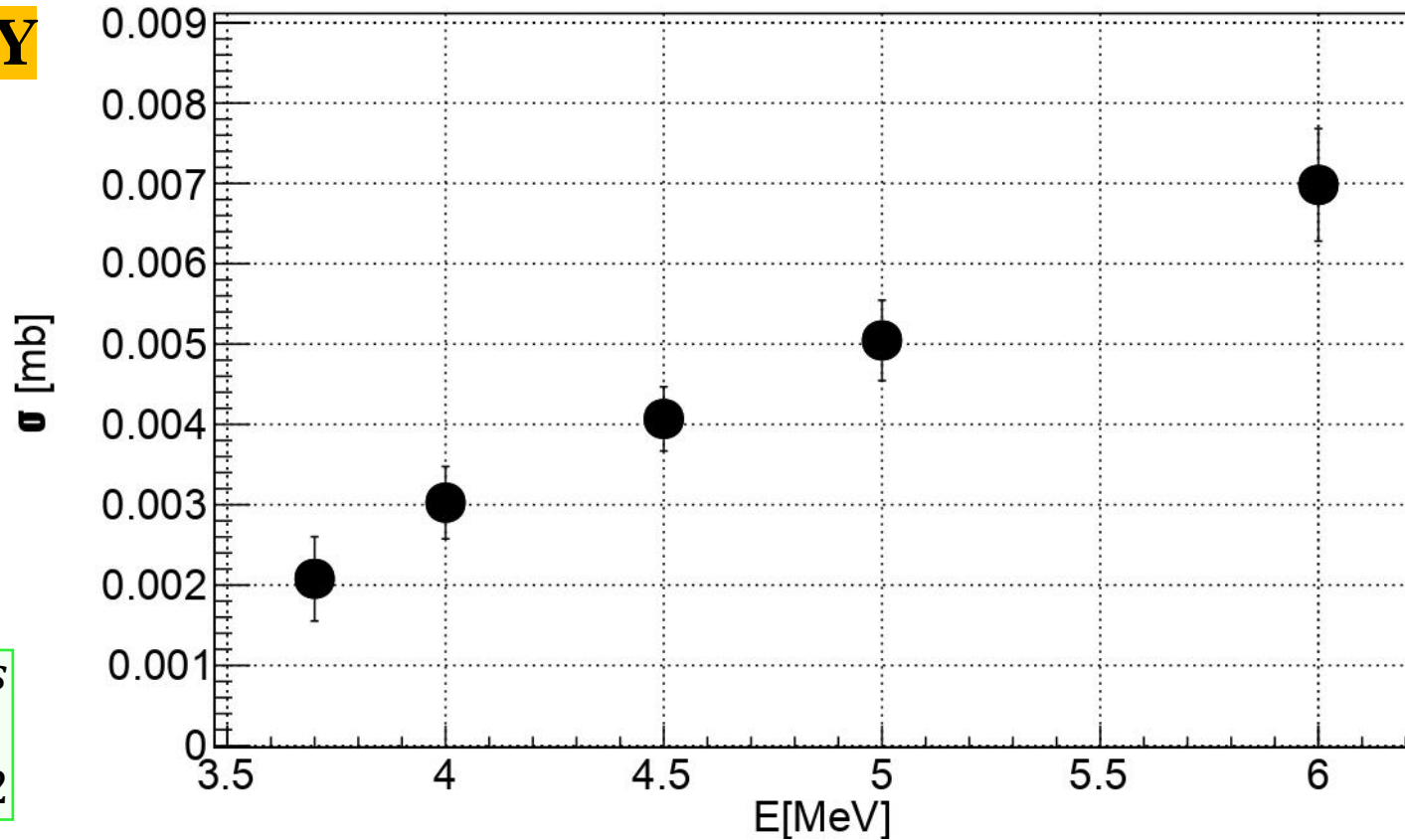


The ground state cross-section of the direct ${}^3\text{H}(\alpha, \gamma){}^7\text{Li}$ reaction

**PRELIMINARY
RESULTS**

Ground state $\Rightarrow j = s$

Photons $\Rightarrow 2j + 1 = 2$



**PRELIMINARY
RESULTS**

$$E_{\alpha t} = E_{\gamma} - Q_{\alpha t \rightarrow {}^7\text{Li}\gamma}$$

$$Q_{\alpha t \rightarrow {}^7\text{Li}\gamma} = 1.397 \text{ MeV}$$

$$\frac{\sigma_{{}^7\text{Li}\gamma \rightarrow \alpha t}}{\sigma_{\alpha t \rightarrow {}^7\text{Li}\gamma}} = \frac{(2j_{\alpha} + 1)(2j_t + 1)}{2 \cdot (2j_{\text{Li}} + 1)} \cdot \frac{2 \cdot m_{\alpha t} \cdot E_{\alpha t} \cdot c^2}{E_{\gamma}^2} \cdot \frac{1}{1 + \delta_{\alpha t}} = \frac{1}{4} \cdot \frac{2 \cdot m_{\alpha t} \cdot E_{\alpha t} \cdot c^2}{E_{\gamma}^2}$$

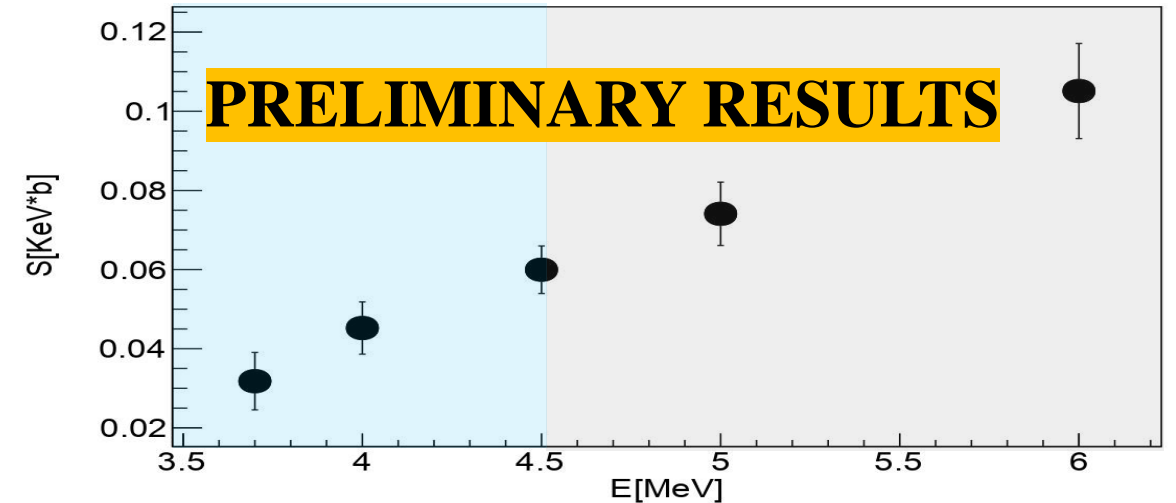
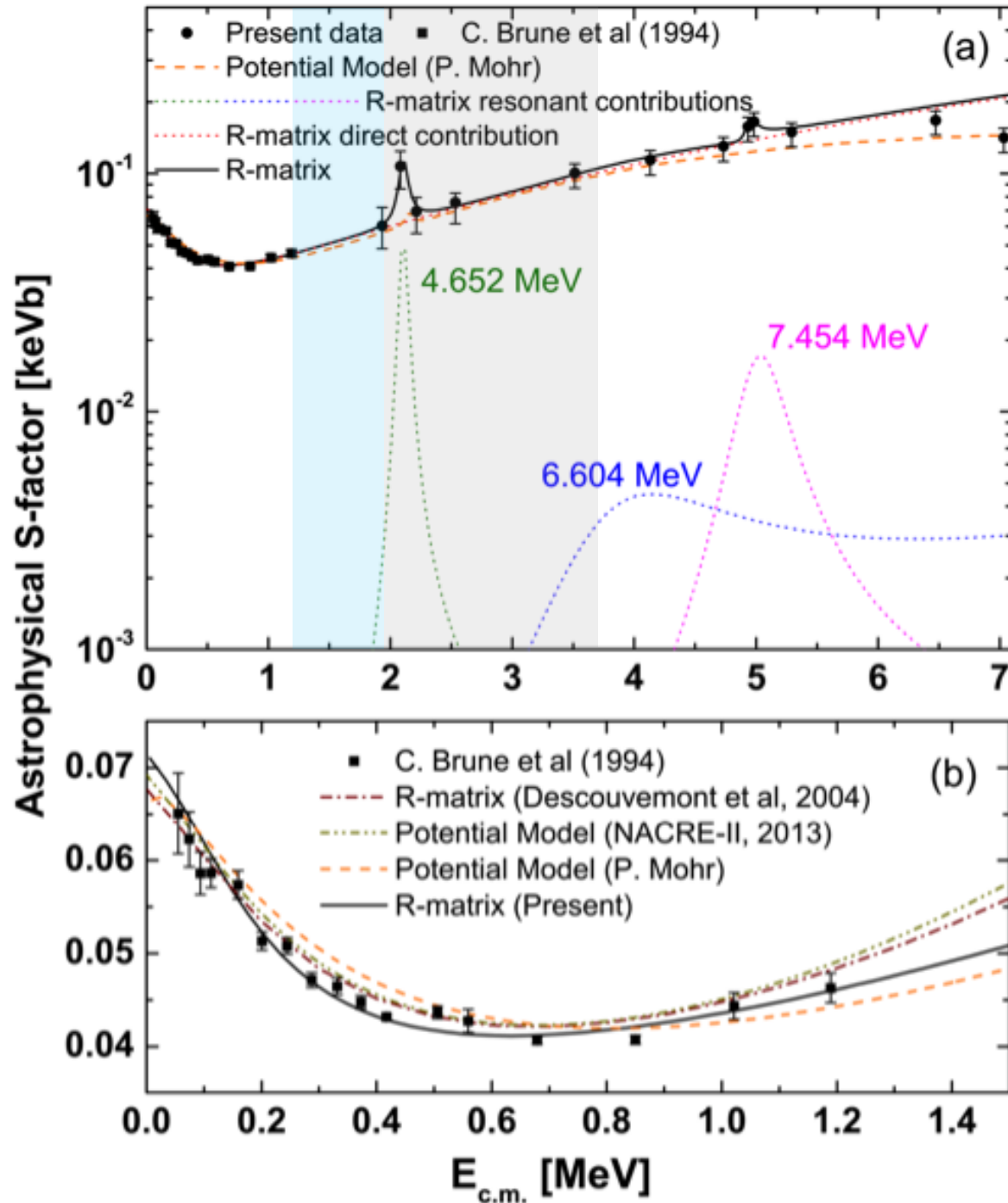
Astrophysical S – factor

$$S_{\alpha t \rightarrow {}^7\text{Li}\gamma}(E) = e^{2\pi\eta} \cdot E \cdot \sigma_{\alpha t \rightarrow {}^7\text{Li}\gamma}(E)$$



$$0.989534 \cdot Z_{\alpha} Z_t \sqrt{\frac{1}{E} \cdot \frac{M_{\alpha} M_t}{M_{\alpha} + M_t}}$$

Measurement of the ${}^7\text{Li}(\gamma, t){}^4\text{He}$ ground-state cross section between $E_{\gamma} = 4.4$ and 10 MeV, M. Munch, C. Matei, S.D. Pain, K.A. Chipps, et al., *Phys. Rev. C* 101, 055801 (2020)



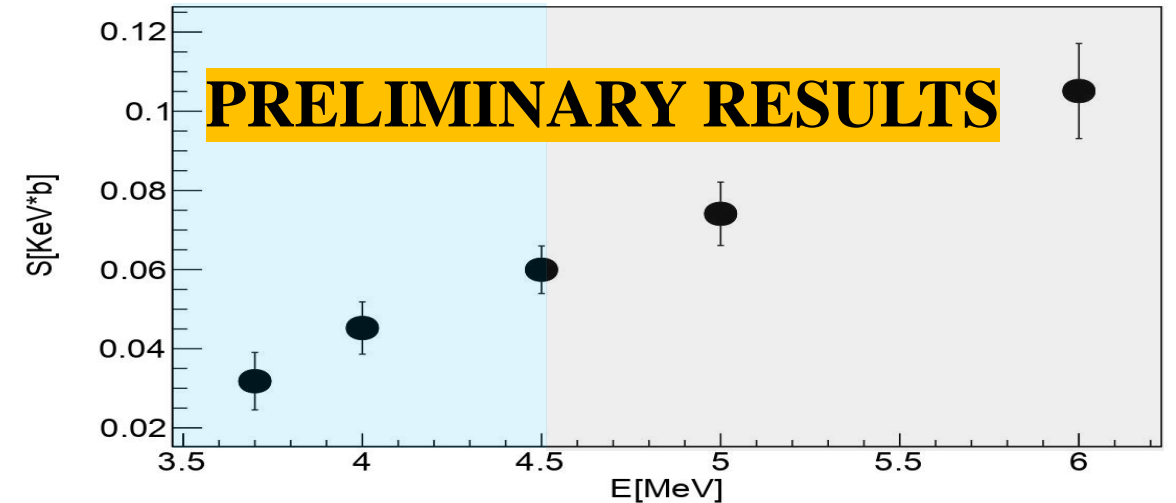
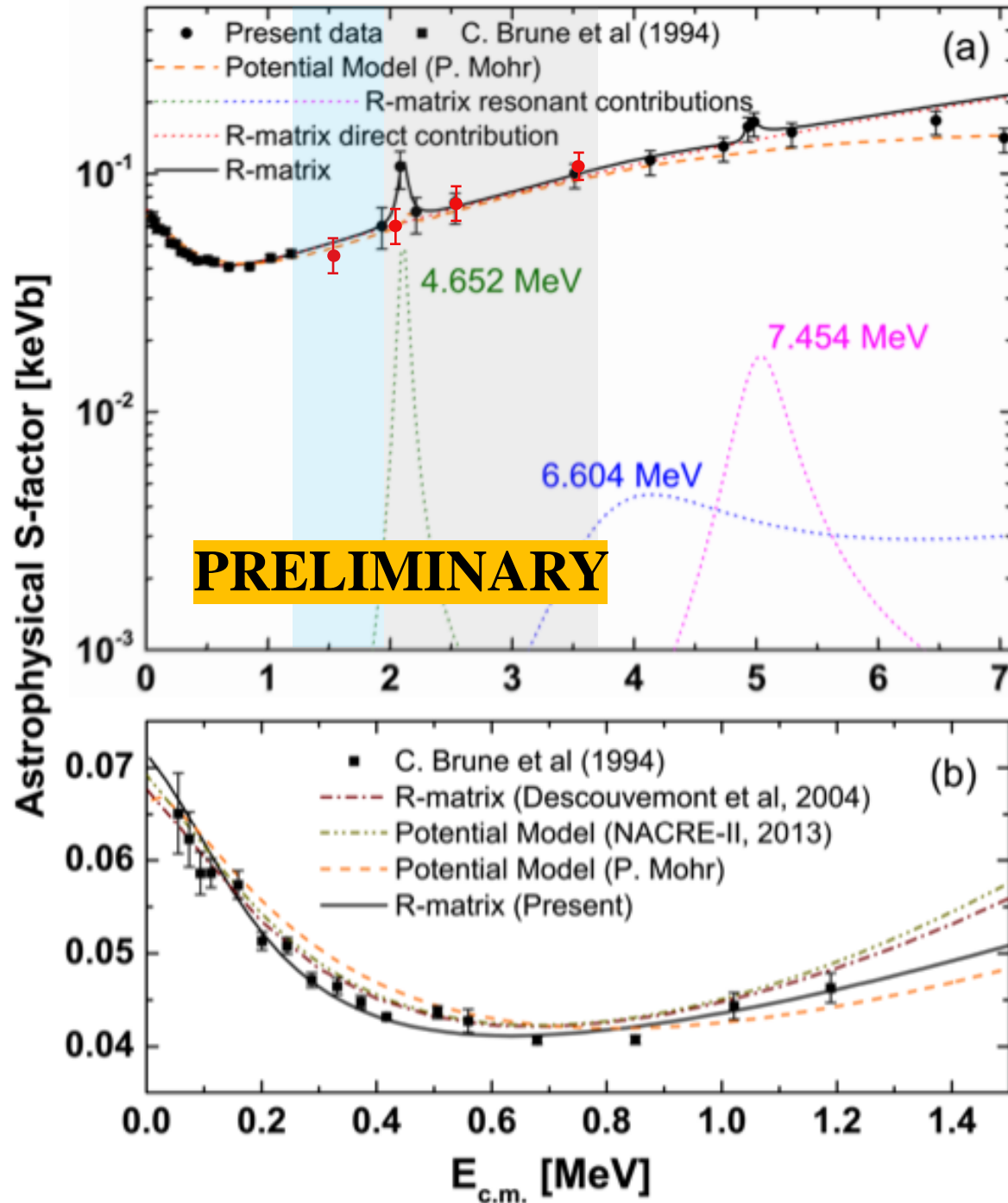
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CONCLUSIONS...

- **The reaction ${}^3\text{H}(\alpha, \gamma){}^7\text{Li}$ contributes to the production of ${}^7\text{Li}$ in Universe** and it's measurement is important due to the discrepancy between the theoretical models and the experimental data.
- **The direct reaction was studied in 1994** by Brune et al for gamma rays energies below 3.7 MeV, but such a measurement cannot be performed anymore.
- **The inverse reaction has been studied in 2017** by our team at HIgS for gamma beam energies between 4.4 and 10 MeV using a Silicon Detector Array. Below 6 MeV the coincidences were clearly observed only in the thinner detectors.
- **A new experiment was performed in April 2023 to cover the gap between 3.7 and 6 MeV.** The set-up was improved by using an array of thinner silicon detectors, by decreasing the thickness and the electron density of the entry flange and by using a longer pipe.
- ✓ **The coincidences have been clearly separated and the background has been highly reduced, affecting only the lowest energy. Data analysis underway. The preliminary ground state cross-section and the s-factor have been successfully extracted.**



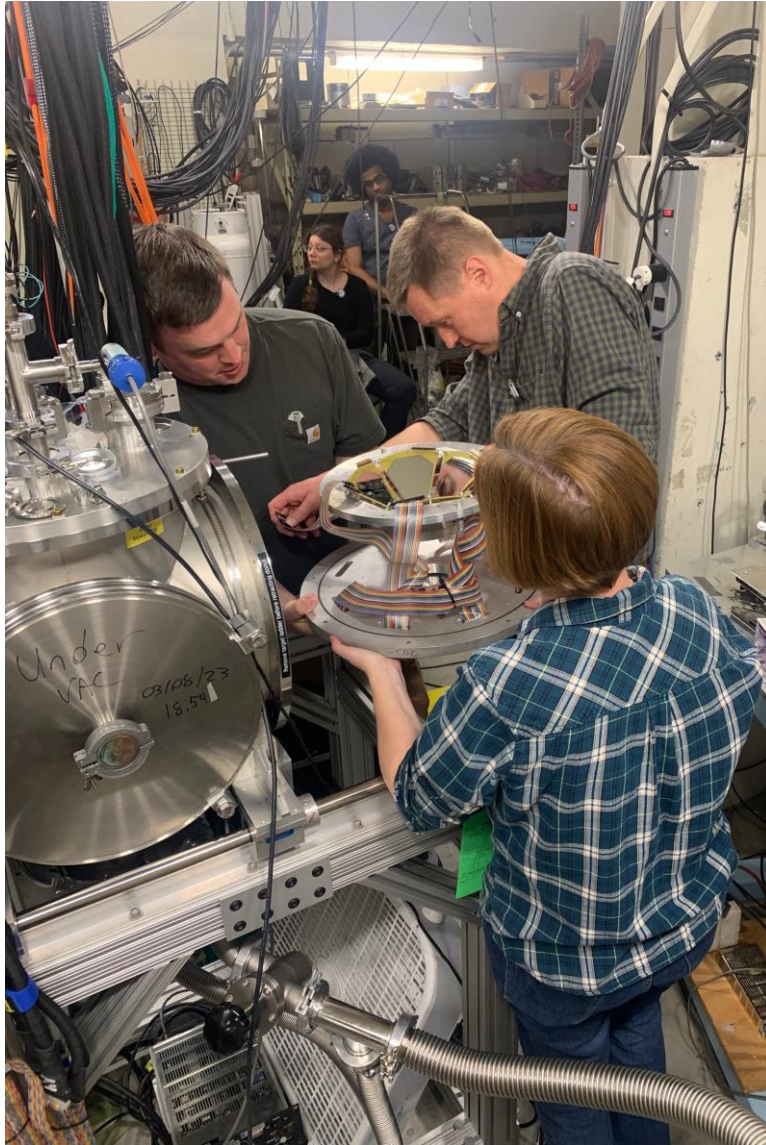
Thank you!



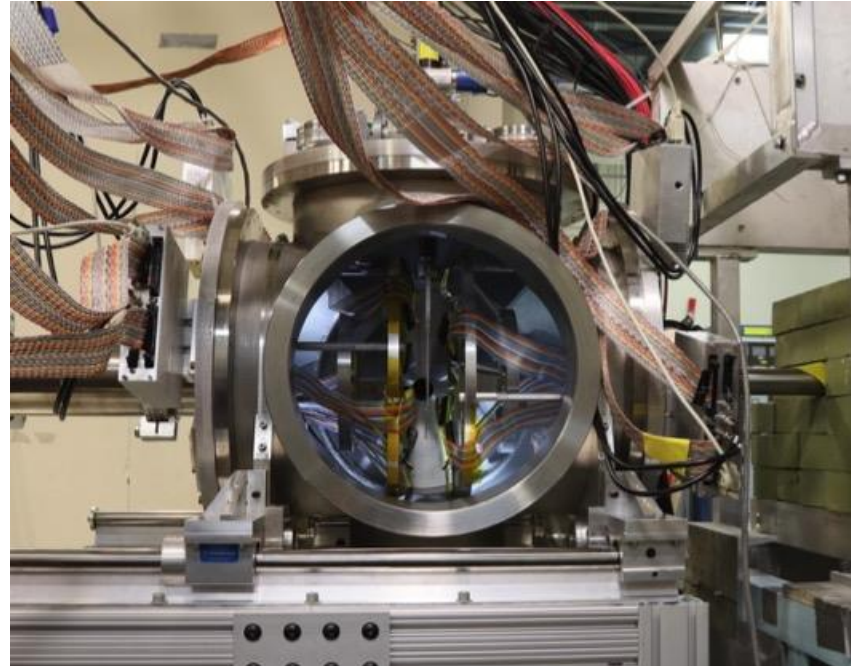
- *Catalin Matei, D. Balabanski, Ioana Kuncser, T. Petruse, A. Pappalardo, H. Pai, Yi Xu (ELI-NP, Bucharest-Magurele, Romania)*
- *Steve Pain, Kelly Chipps, T. King, M Febbraro (ORNL – Oak Ridge, Tennessee, USA)*
- *M. Grinder, S. Balakrishnan, Heather Garland, Jolie Cizewski (Rutger University - New Brunswick, New Jersey, USA)*
- *H. Karwowski, R.V. Janssens, T. Psaltis, C. Marshall (University of Nord Carolina - Chapel Hill, Nord Carolina USA)*
- *C.R. Brune, A. Voinov (Ohio University, Athens, Ohio, USA)*
- *O. Tindle, C. Haverson, R. Smith (Sheffield Hallam University, UK)*
- *M. La Cognata, G.L. Guardo, S. Palmerini, L. Sergi, D. Lattuada, R.G. Pizzone, G. Rapisarda, A. Tumino (INFN-Laboratori Nazionali del Sud, Catania, Italy)*
- *K.Y. Chae, Gyoungmo GU (Sungkyunkwan University, Suwon, Korea)*

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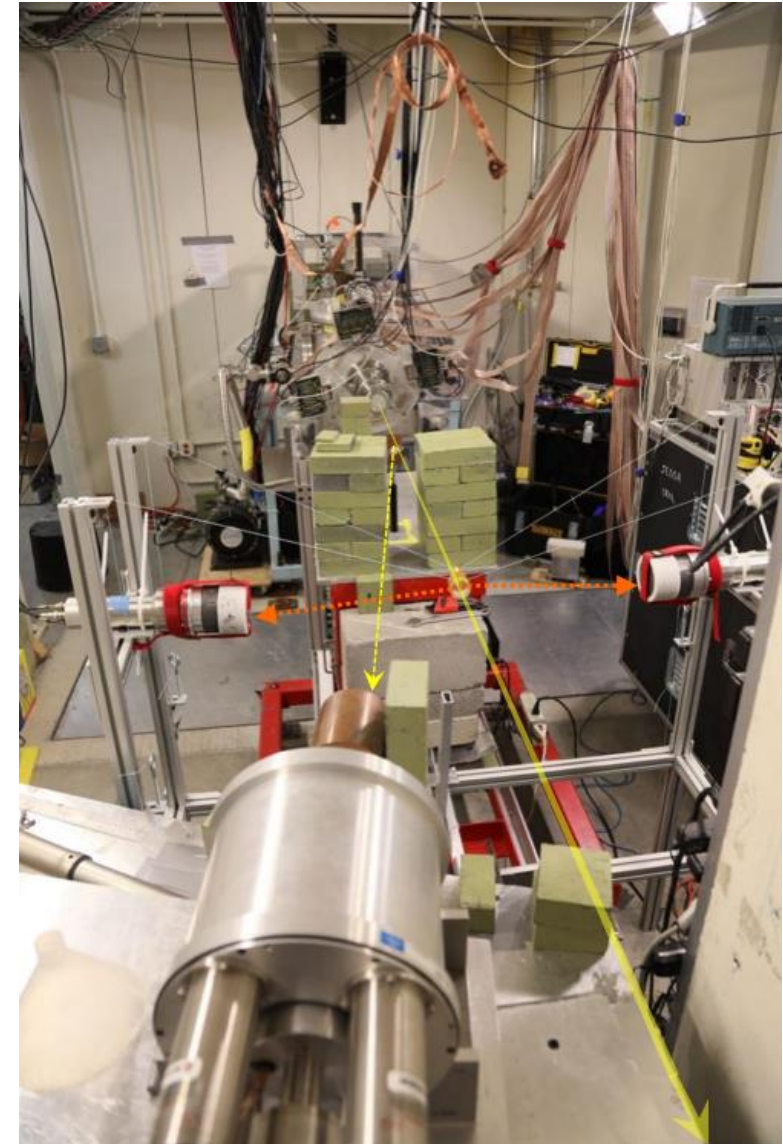
Set-up at H_γS



Mounting detectors

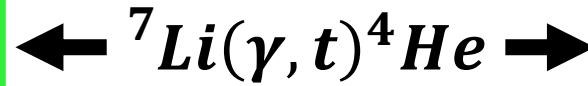


Vacuum chamber for charged-particle detection



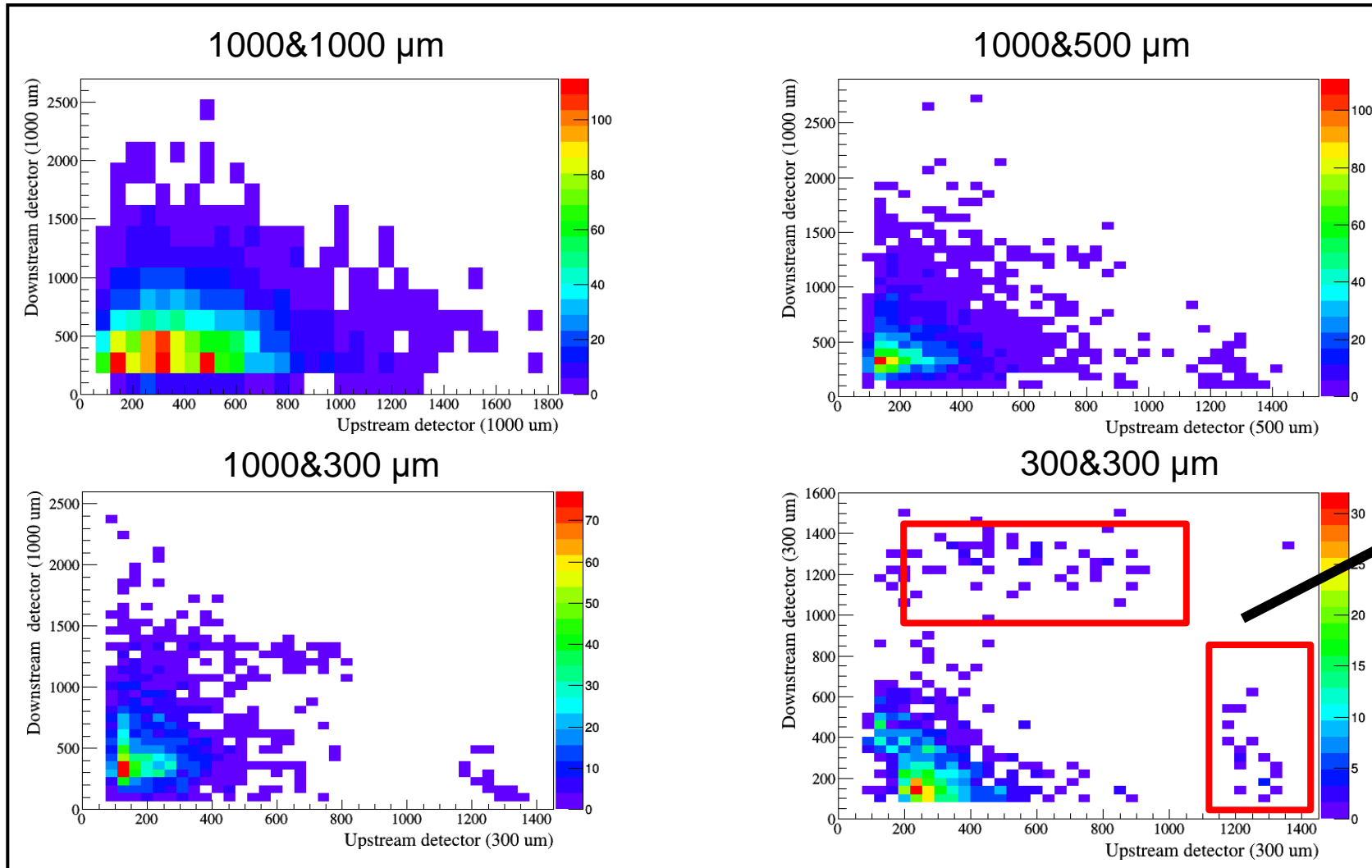
Beam diagnostics

The old experiment:
2017
 4.4 MeV - 10 MeV



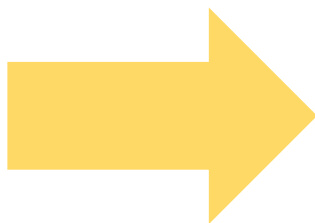
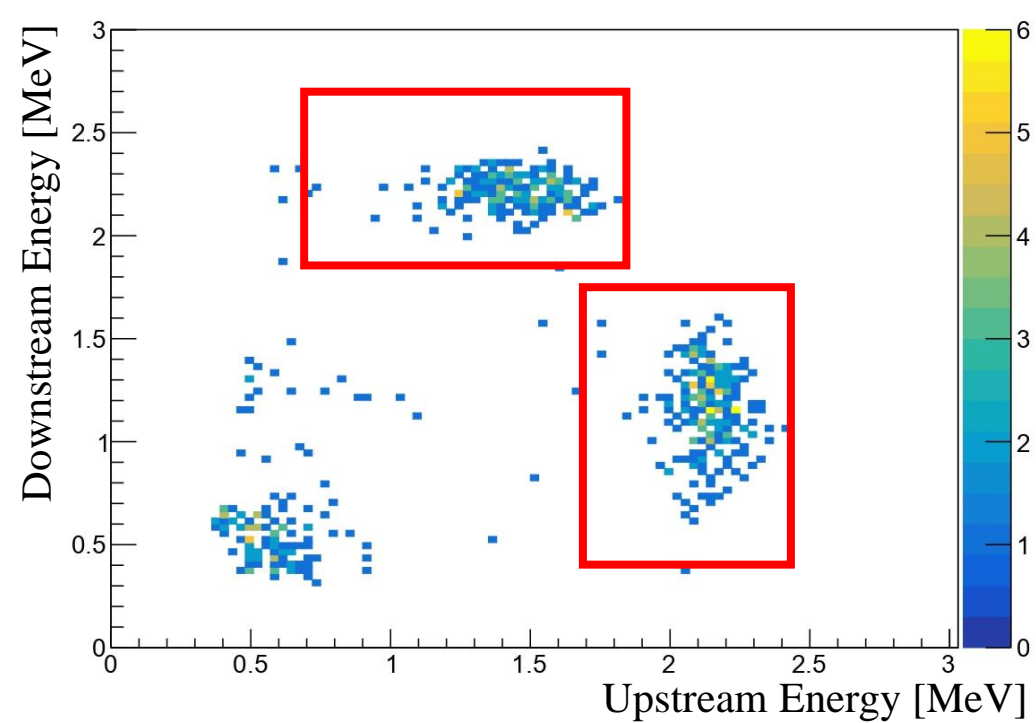
The new experiment:
6th April 2023-12th April 2023
 3.7 MeV - 6 MeV

HIγS
Laboratory
of
Duke
University
(USA),
2017

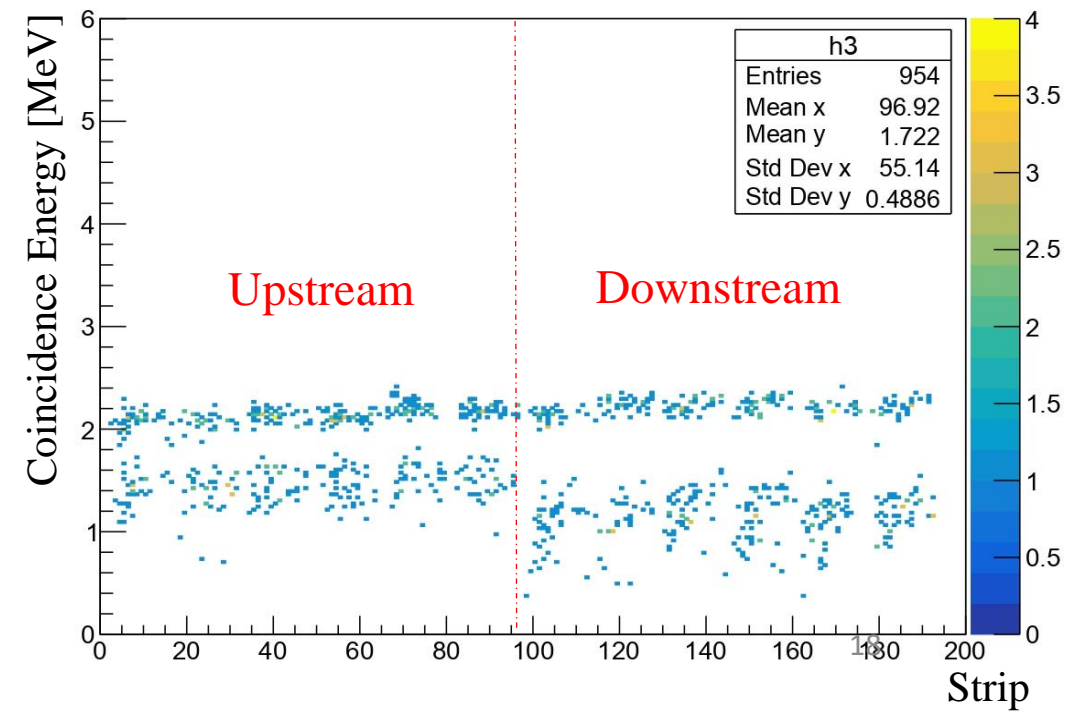
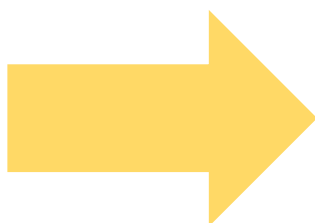
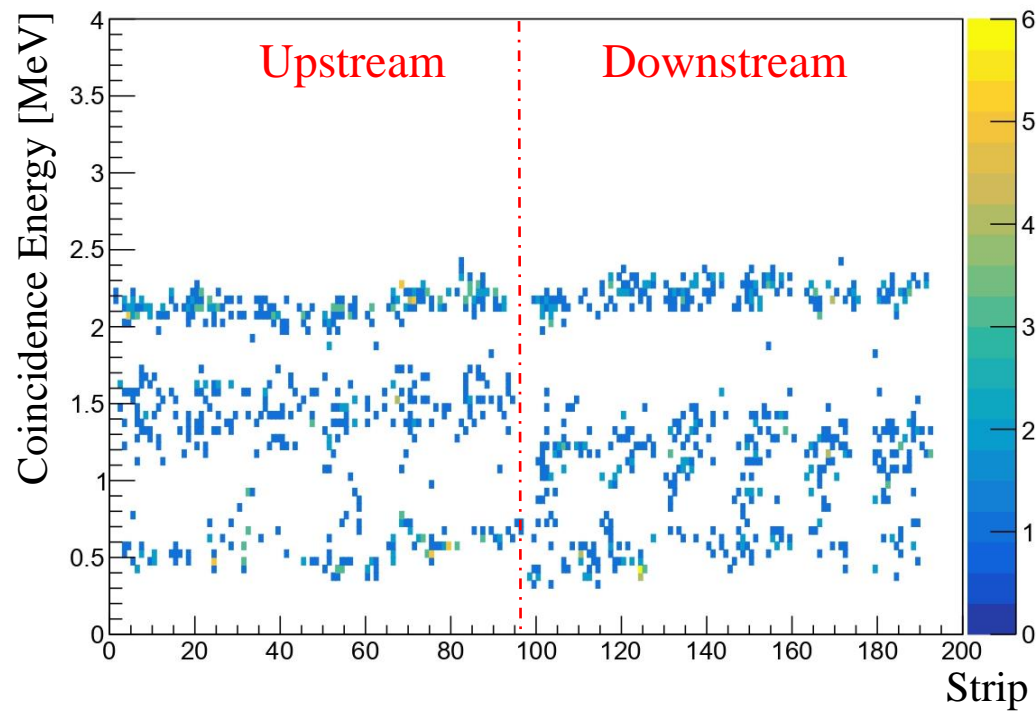
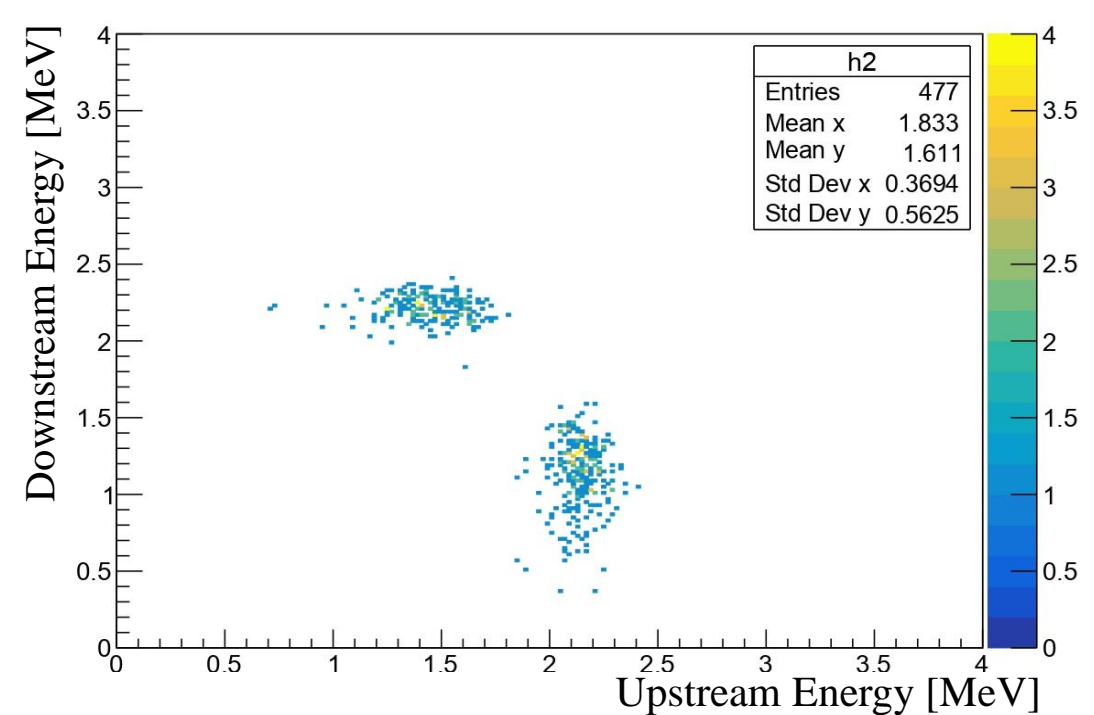
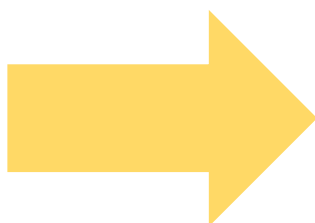


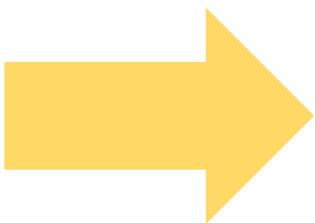
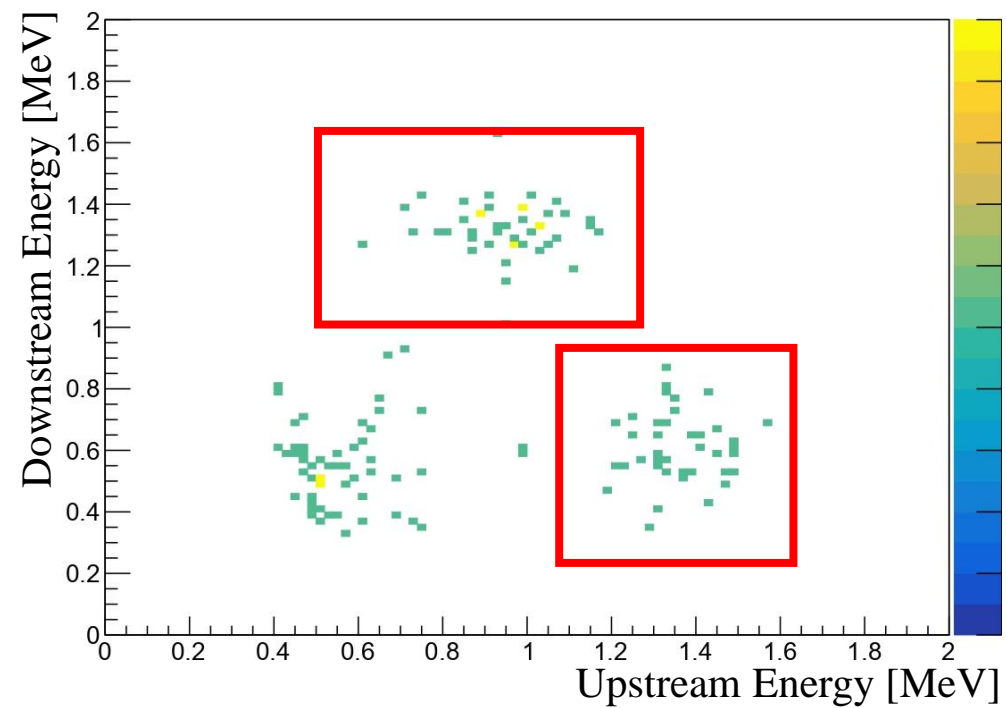
The
coincidences
were
clearly
separated for
energies
higher than 6
MeV

Coincidences
 corresponding
 to 4.4 MeV
 gamma beam
 were observed
 only in the
 thinner
 detectors

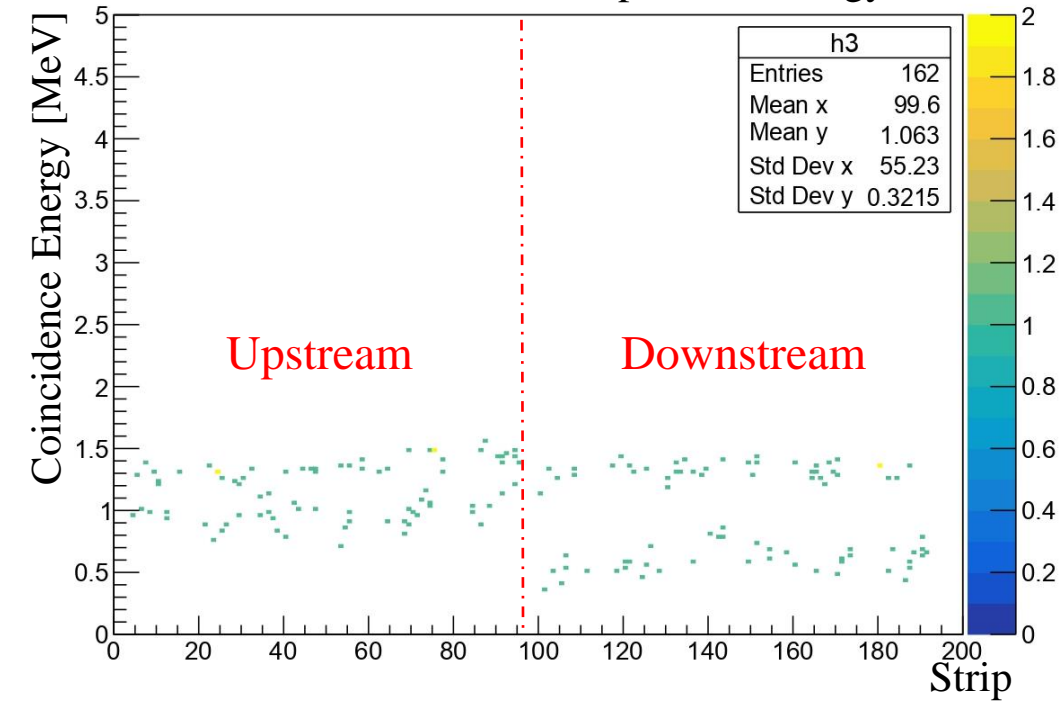
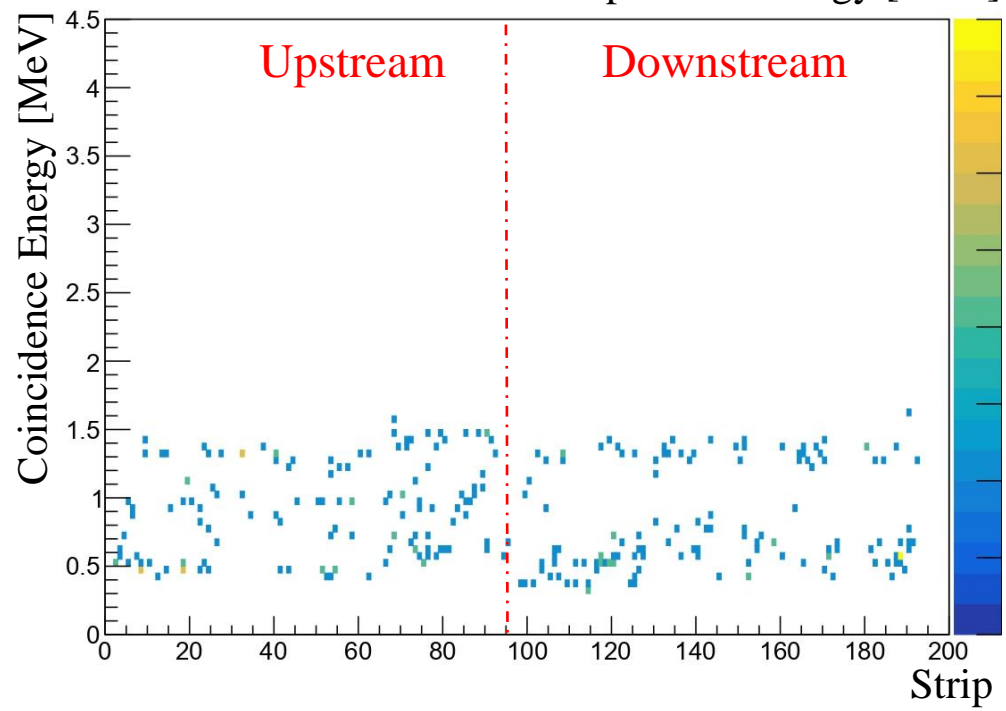
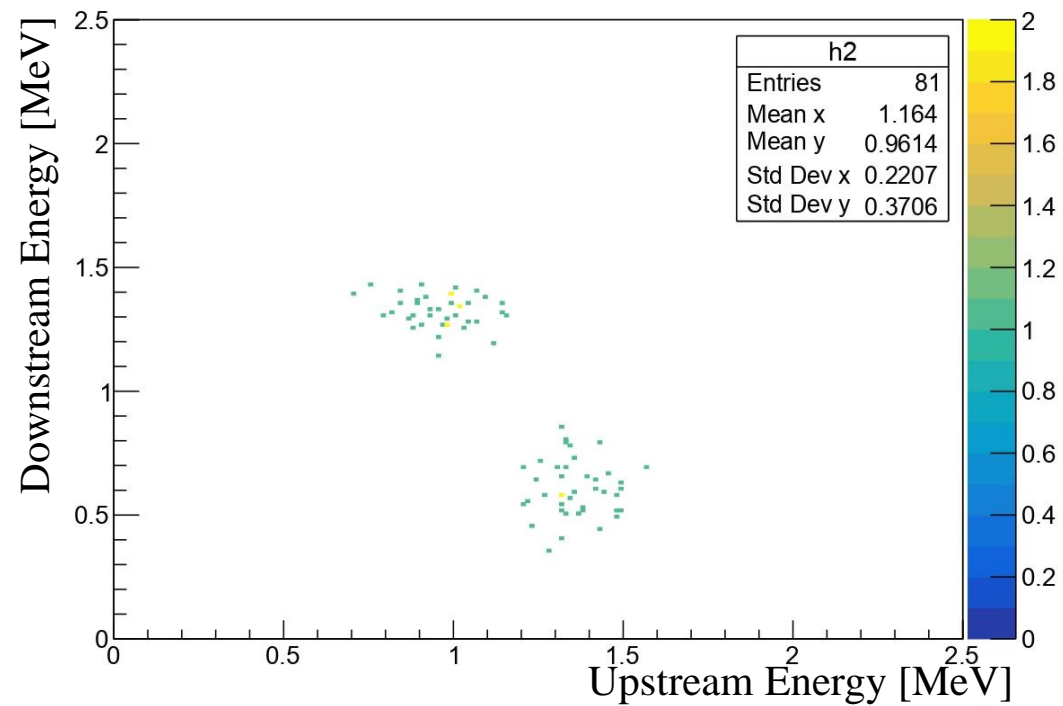
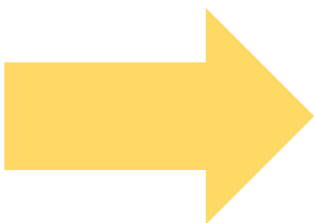


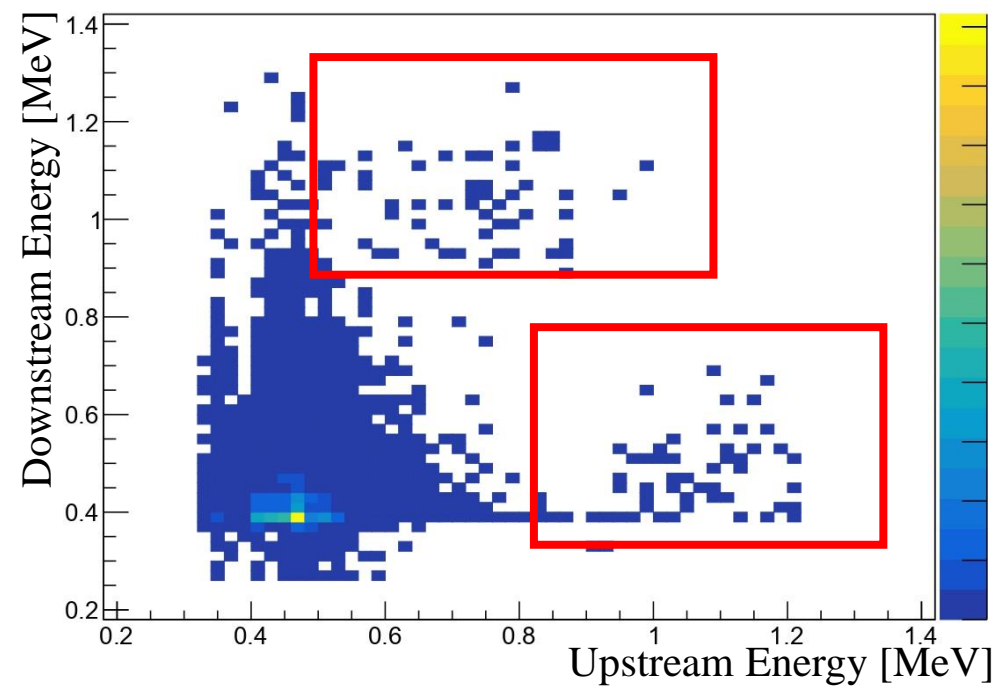
Background
can be easily
subtracted
for
6 MeV



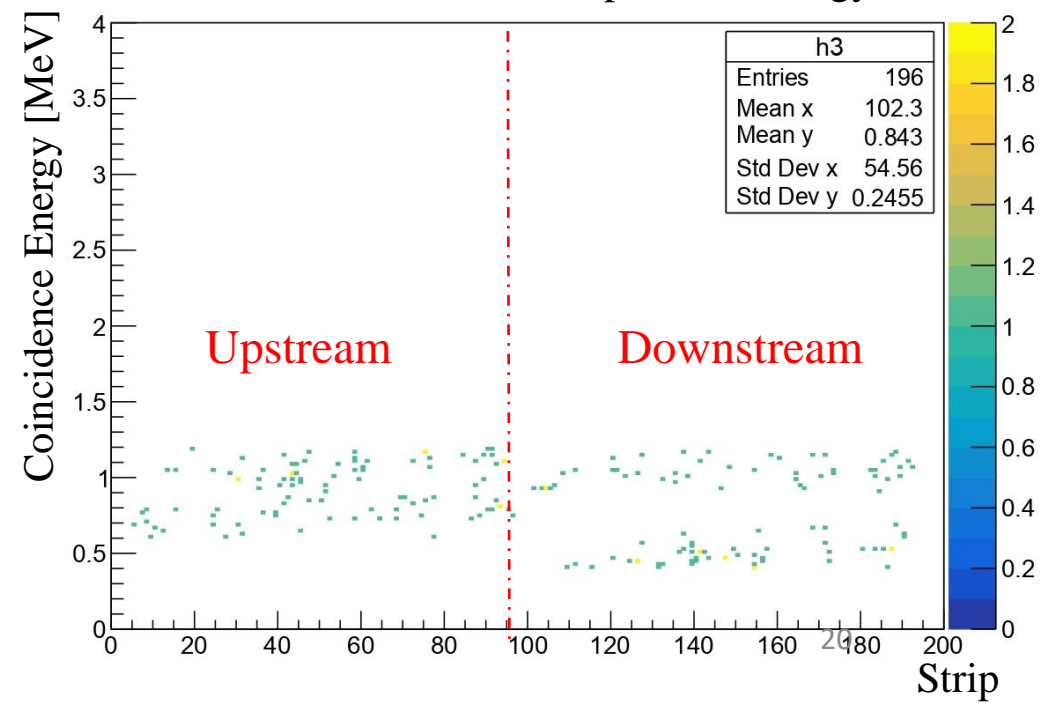
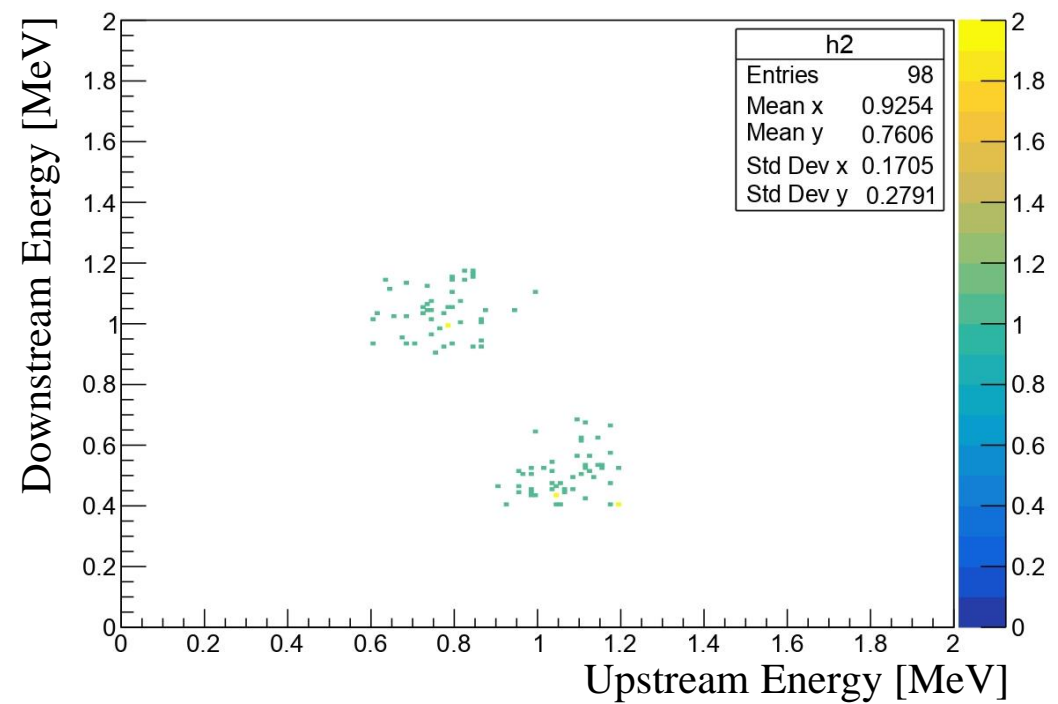
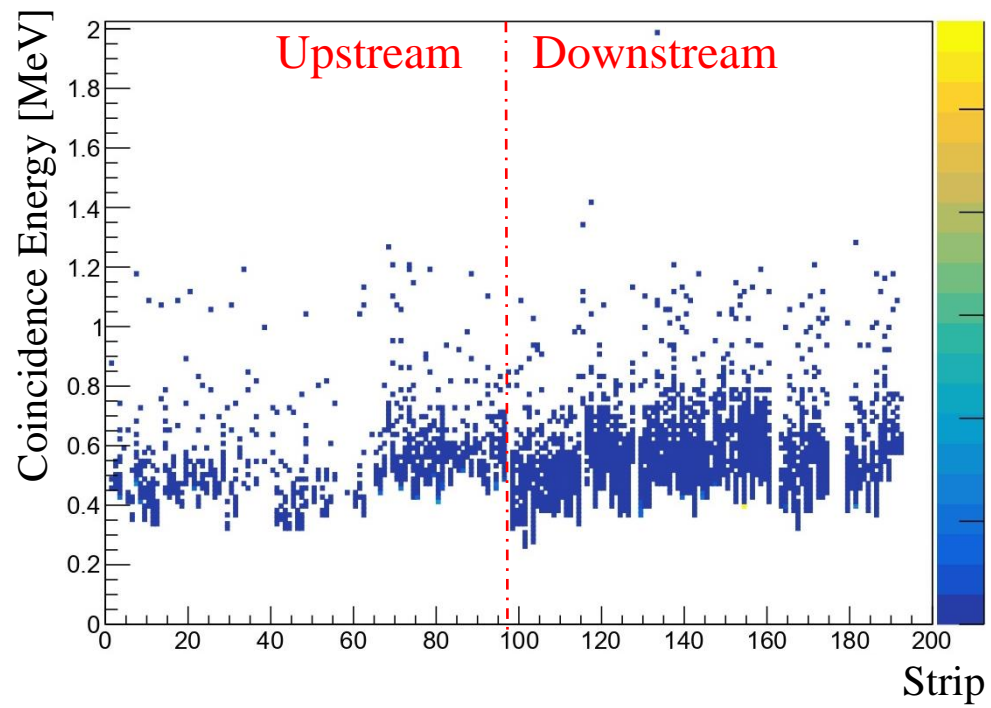
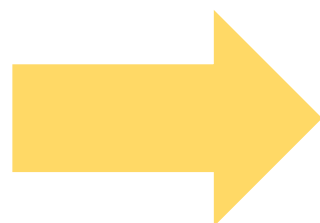


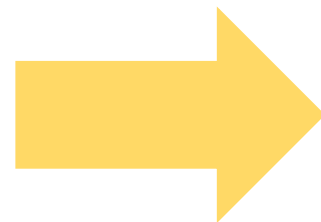
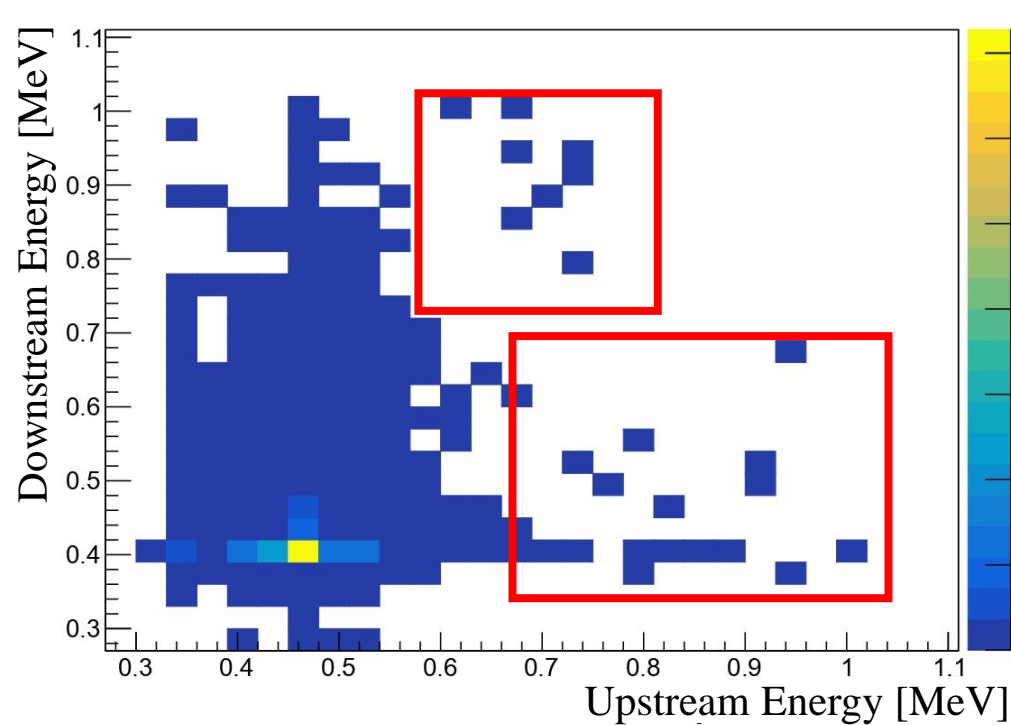
And
even
for
4.5 MeV



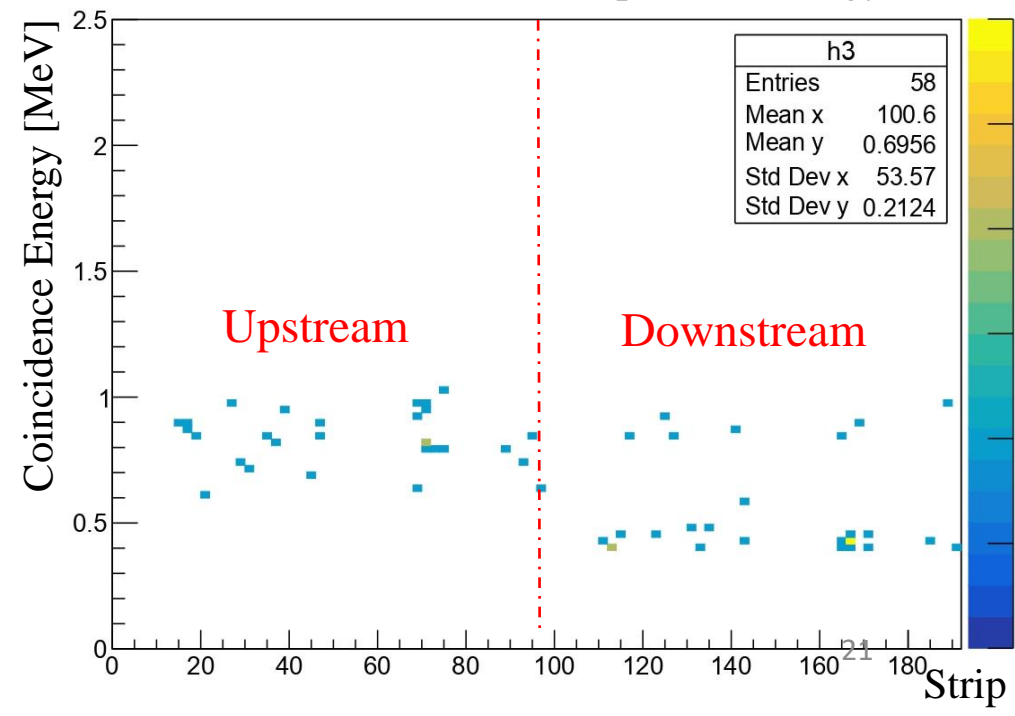
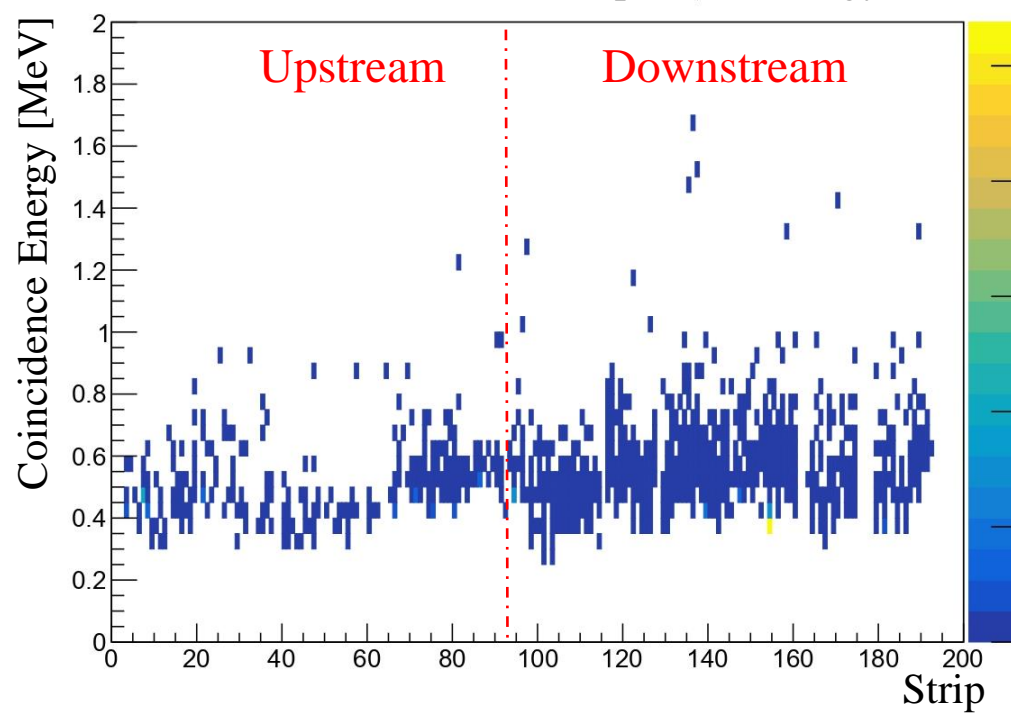
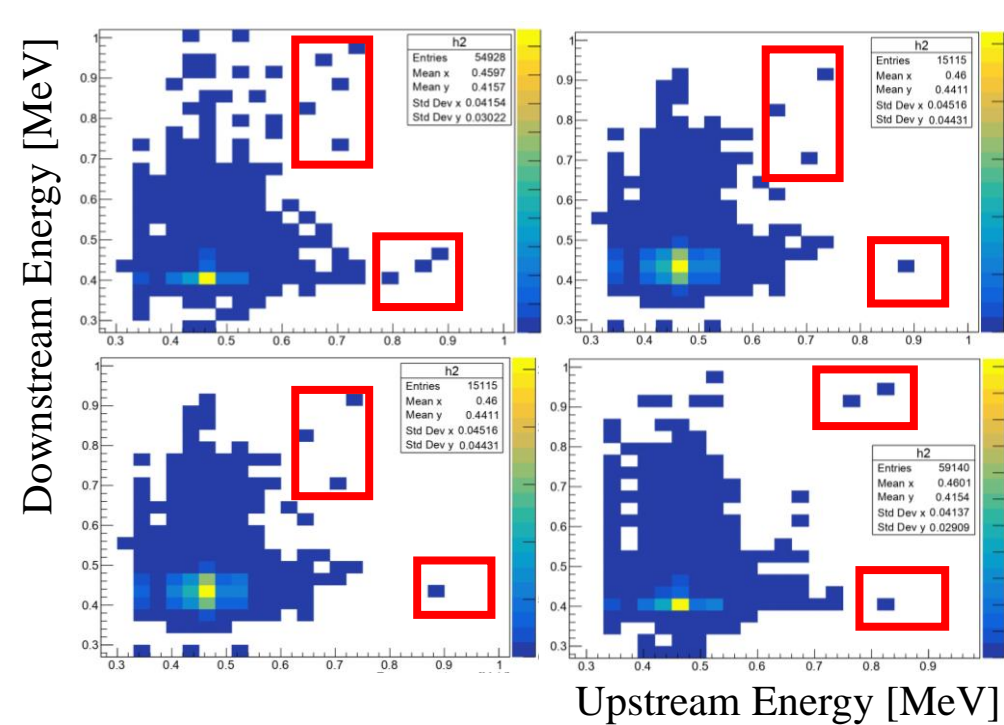
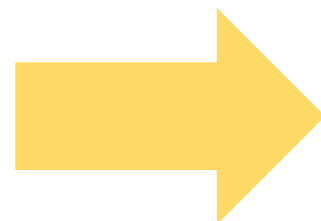


Still
possible
for
4 MeV

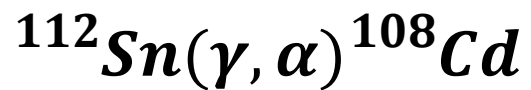
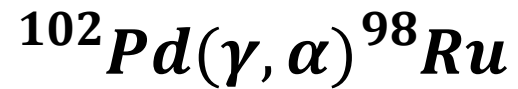
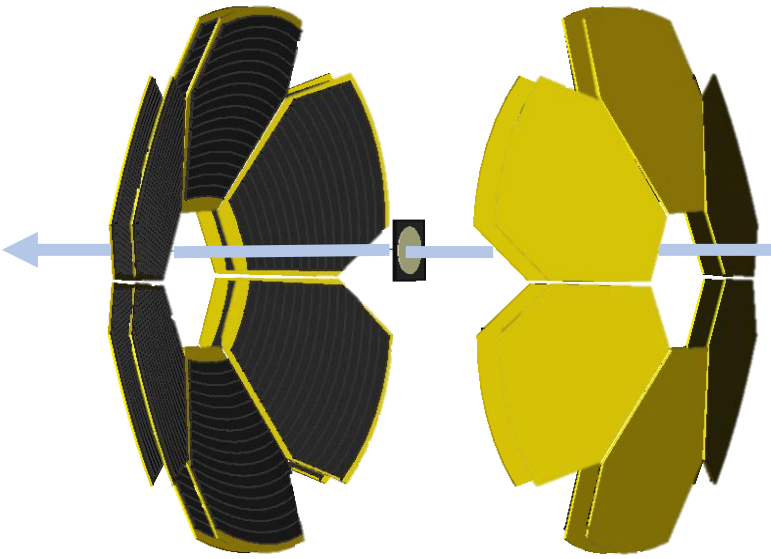




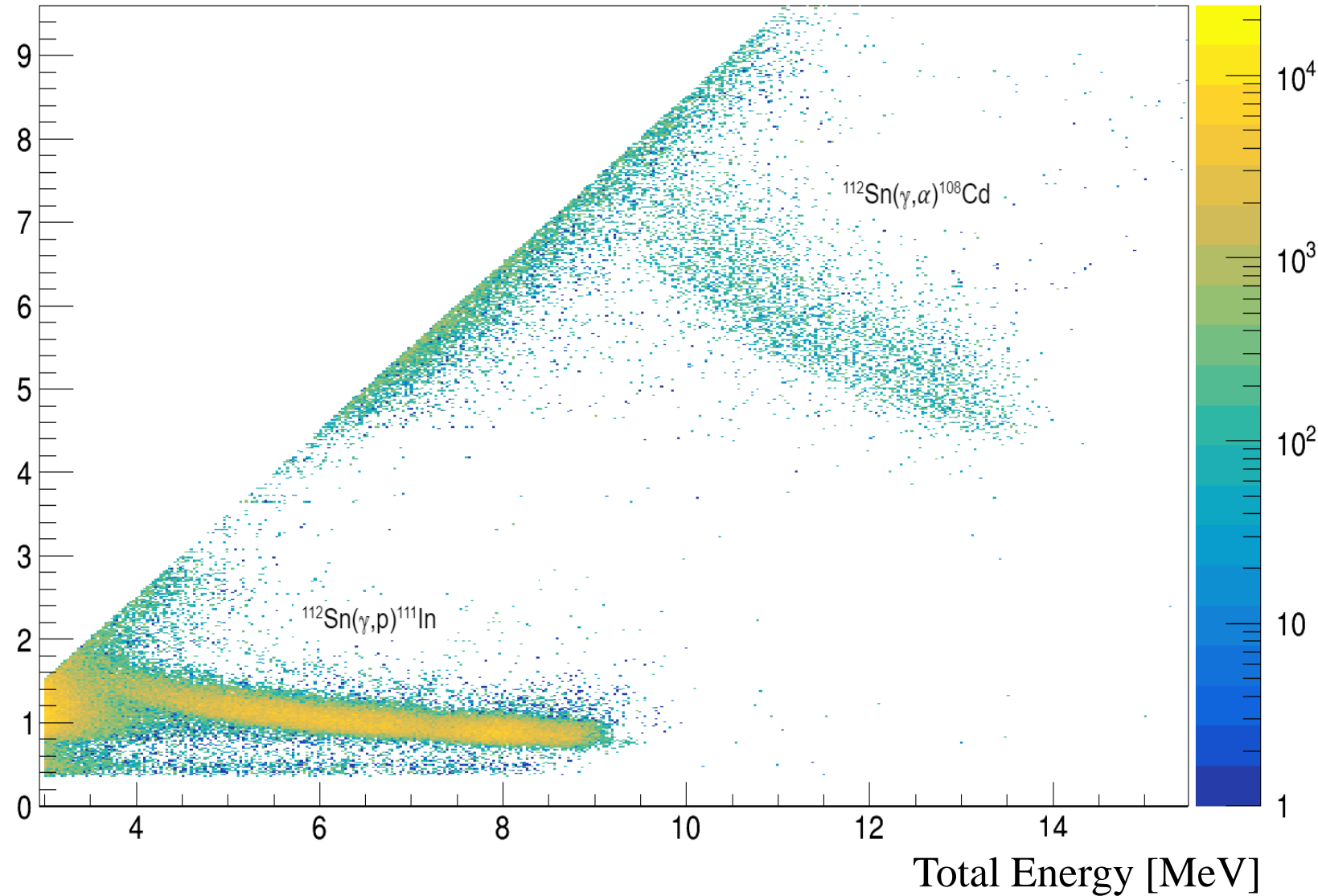
Difficult to distinguish
for
3.7 MeV



p – process



ΔE [MeV]



Gamma beam energy measurement

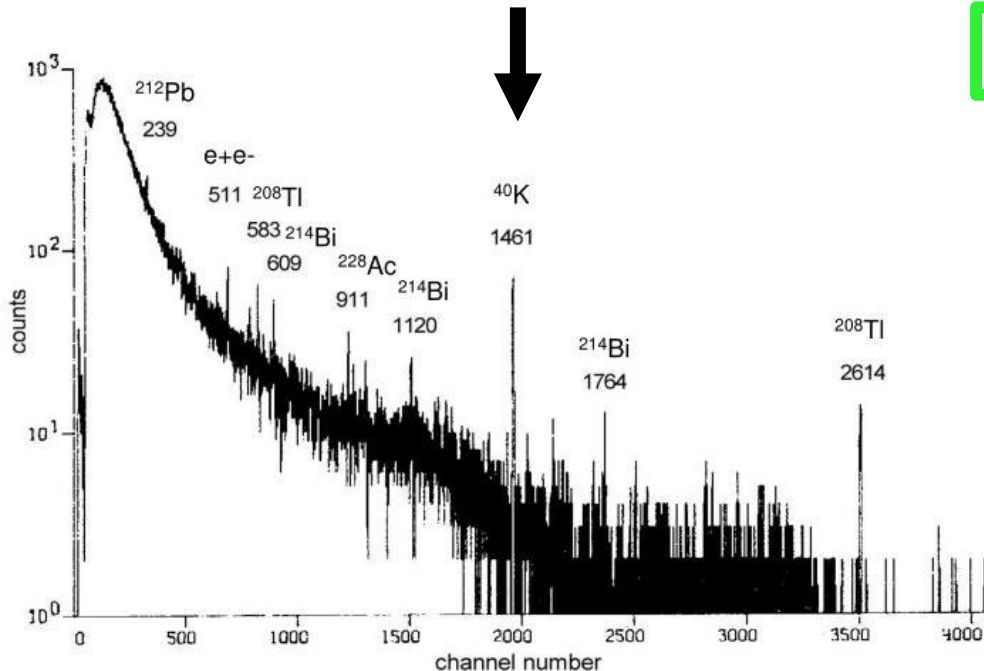
Ge DETECTOR + ATTENUATORS

Calibrated using background

Known energy loss

$$E_{\gamma} = E_{det} + E_{loss}$$

39.35 cm
of
copper

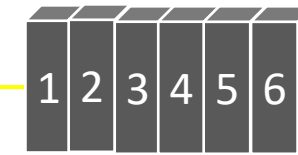
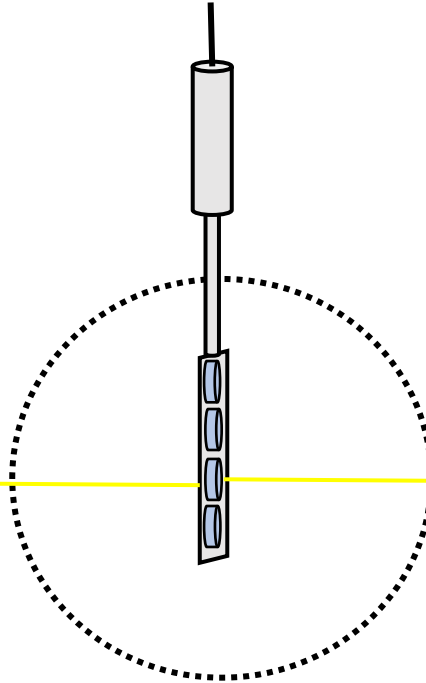
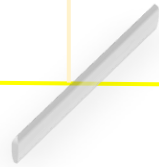
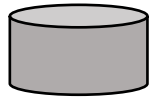


Gamma beam intensity measurement

Measuring
continuously

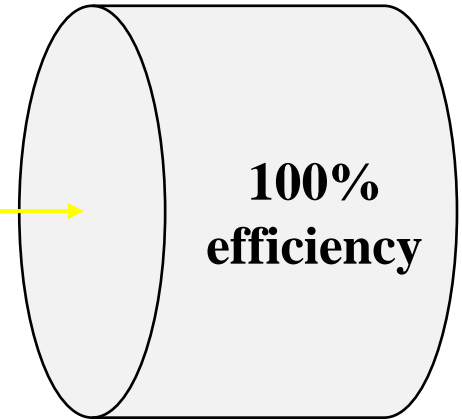
Pb glass

gamma beam

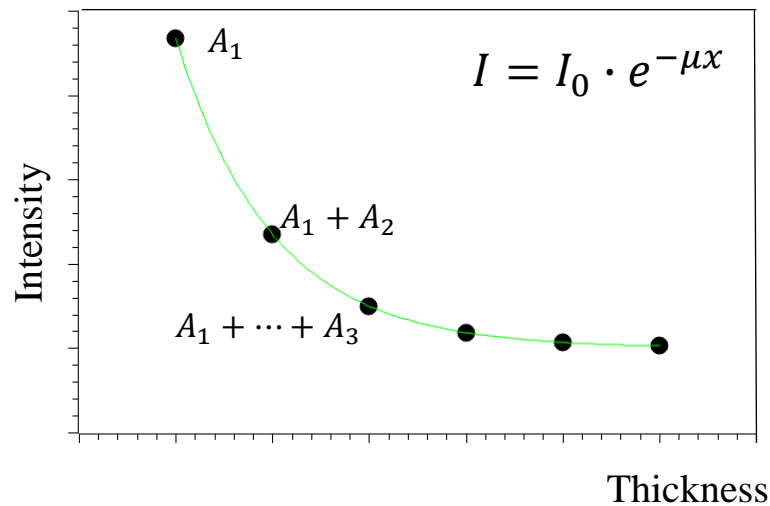


Attenuators

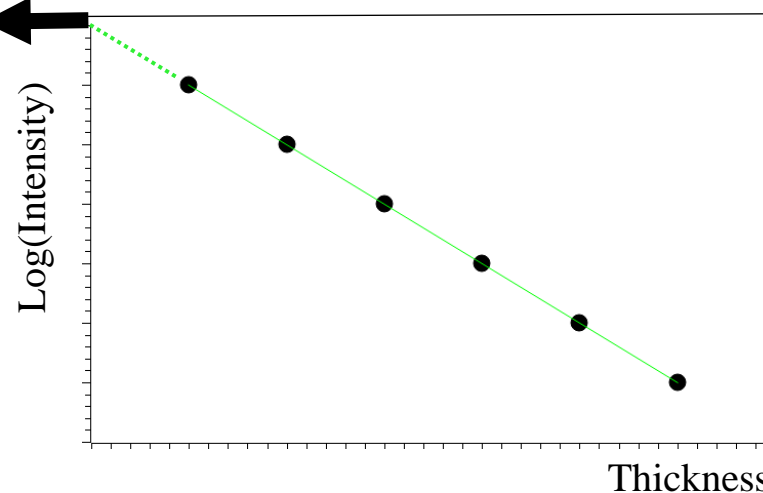
NaI Detector



**100%
efficiency**



No
Attenuation
Value

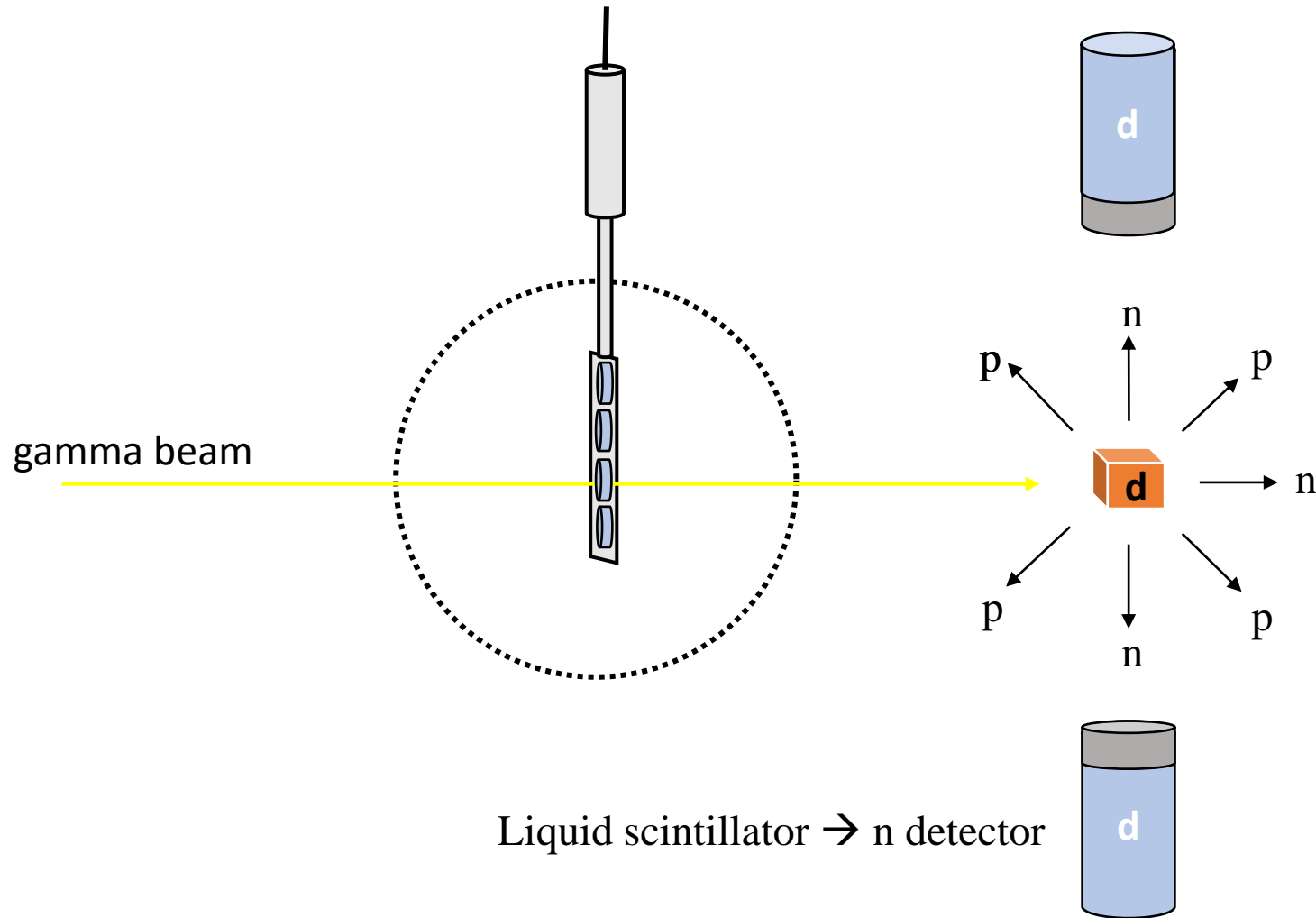


*Coefficient related to the
energy of the beam*

Pb glass intensity calibration

Intensity measurement during
the run

Gamma beam intensity measurement



deuterium photodisintegration

$$d + \gamma = n + p + 2.2 \text{ MeV}$$



$$E_n = \frac{E_{\text{gamma}} - 2.2 \text{ MeV}}{2}$$

$$\begin{aligned} \text{Nr of neutrons:} \\ \varepsilon n = \sigma_{\text{photodis}} \cdot N_d \cdot I_\gamma \end{aligned}$$

