



Extreme Light Infrastructure - Nuclear Physics

Gamma Driven Experiments Department

$p + {}^{11}\text{B}$ fusion measurement with ELISSA

Kabita Kundalia

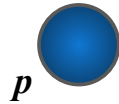
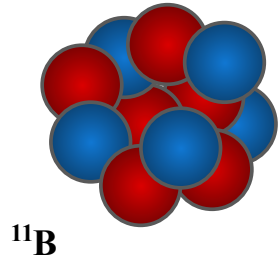
7 - 11th September, 2026

**12th International Conference on Nuclear
Physics in Astrophysics**



Outline :

- ❑ Physics Aspect
- ❑ Experimental Setup
- ❑ Analysis
- ❑ Outlook

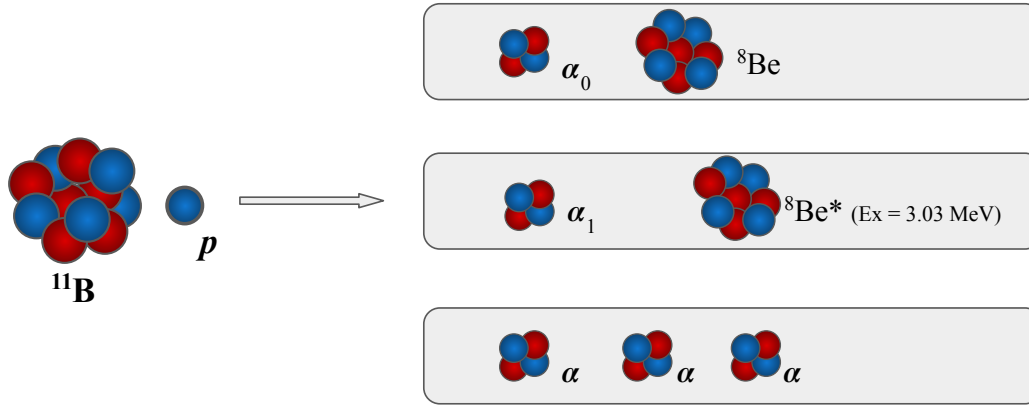


Nuclear Astrophysics:

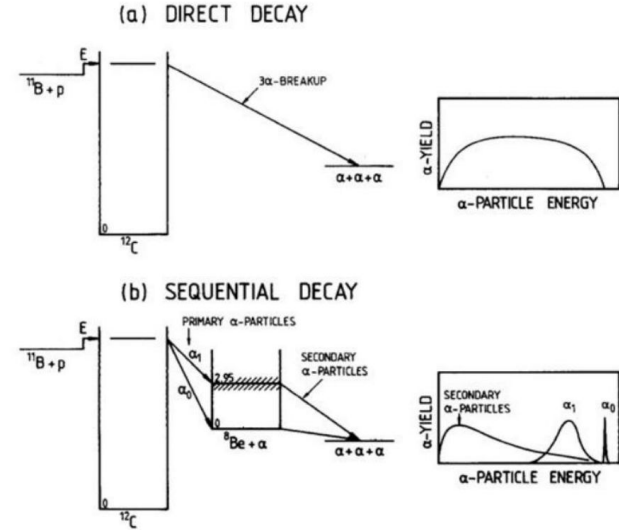
- Study of LiBeB abundances in young main sequence F and G stars can give useful information on **stellar mixing processes** not fully understood yet.
- $^{11}\text{B}(p,\alpha)\alpha$ is the primary mechanism for **^{11}B depletion** within stellar environment.
- To understand if the current discrepancies are only due to a not well-established **theoretical description** of the stellar mixing phenomena, (p, α) nuclear reaction cross sections must be carefully evaluated.

Fusion reactor development:

- $^{11}\text{B}(p,\alpha)\alpha$ reaction has gained prominence primarily due to its relevance in nuclear **fusion reactor** development.
- holds promise for clean energy; reactants, p and ^{11}B , are **abundant, safe**.
- ^{11}B - p fusion **simplifies reactor engineering** compared to DT fusion; avoids challenges associated with high neutron flux and material handling.



Q-value = 8.59 MeV



The proton–boron fusion reaction is dominated by the **two-step sequential decay** via the short-living ${}^8\text{Be}$ in its fundamental or excited state. The direct decay probability is 2000–3000 times less than the sequential decay probability [*Taskaev 2024*].

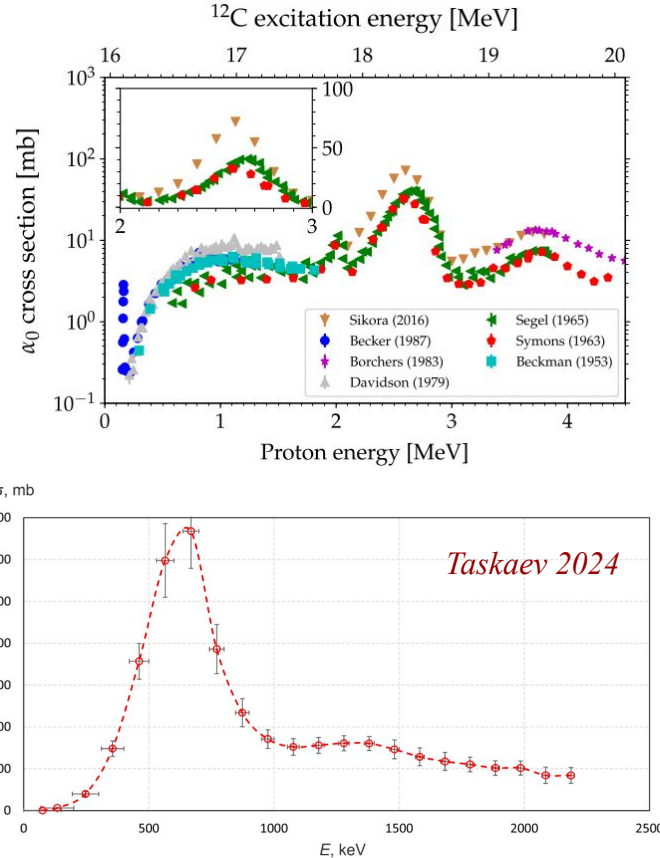


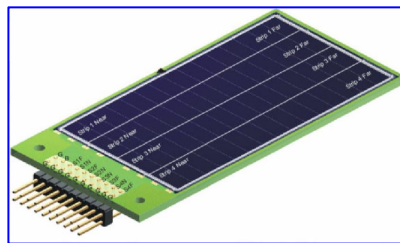
Fig. 13. The measured cross-section of the $^{11}\text{B}(p, \alpha_1)^8\text{Be}^*$ reaction.

There is disagreement in the literature on the interpretation of the cross-section and Legendre coefficients below 2 MeV in terms of resonances in ^{12}C .

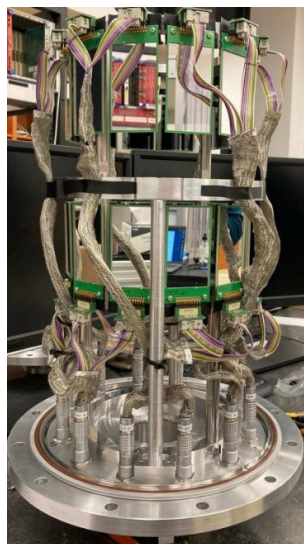
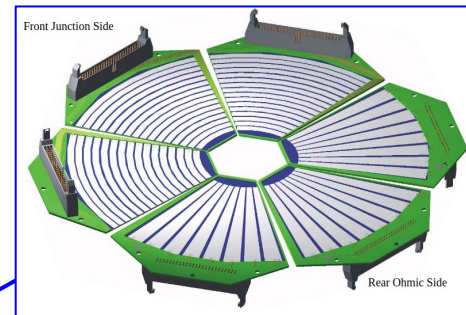
This could not be resolved with a measurement of the α_0 channel. Instead this issue requires a measurement of the α_1 channel and also of the phase-space (Dalitz) distributions of the 3α emission of that channel [*Munch 2020*]. That will provide additional information on the spin and parity of the involved resonances in ^{12}C .

ELISSA

20 Micron X3 PSD, 1000 μm , 160 channels of electronics
~1% energy resolution, 0.3 mm position resolution
41° to 68° and 101° to 133° (lab frame) angular coverage

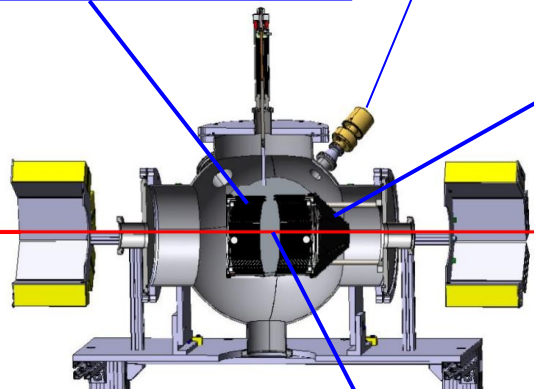


Monitor



Beam

proton - 1 nA
300 keV to 2600 keV



MELISSA

4 Micron MMM DSSSD, 300 μm thick
96 channels of electronics
130° to 175° (lab frame) angular coverage

Target

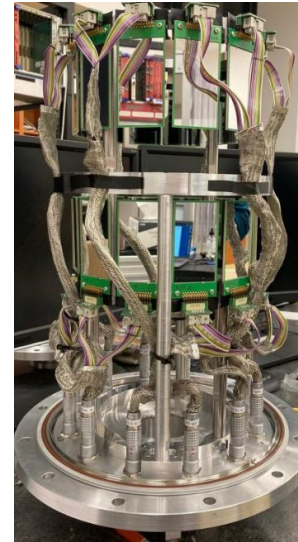
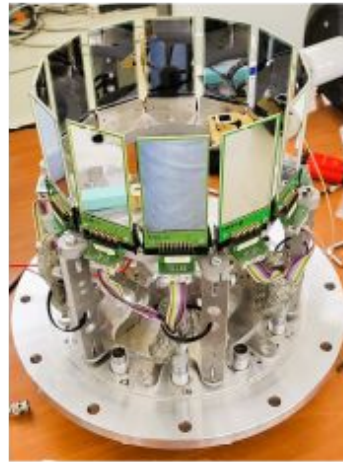
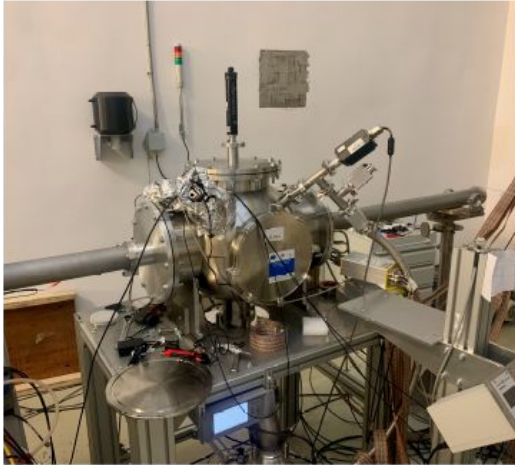
^{11}B ~ 50 - 150 $\mu\text{g}/\text{cm}^2$



In-beam characterization of ELISSA: The 4π position-sensitive silicon strip array at ELI-NP

T. Petruse : published in Nuclear Instrumentation and Methods In Physics Research A

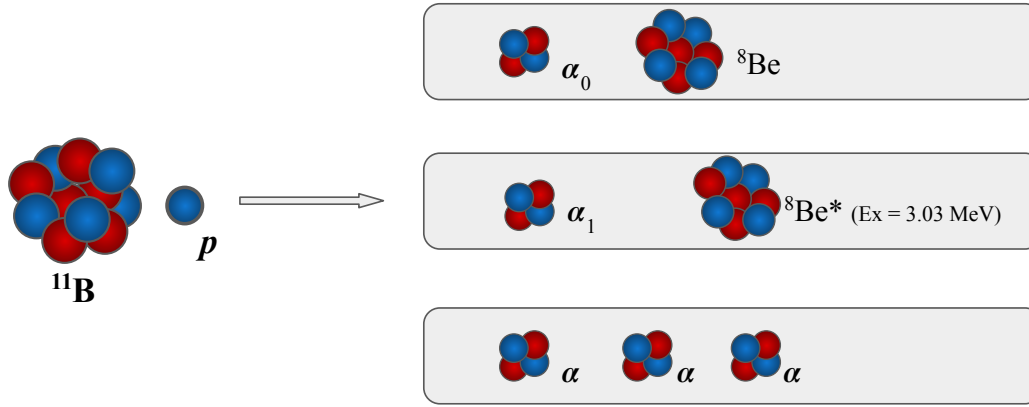
Built and being developed
in collaboration of :
ELI-NP & INFN-LNS



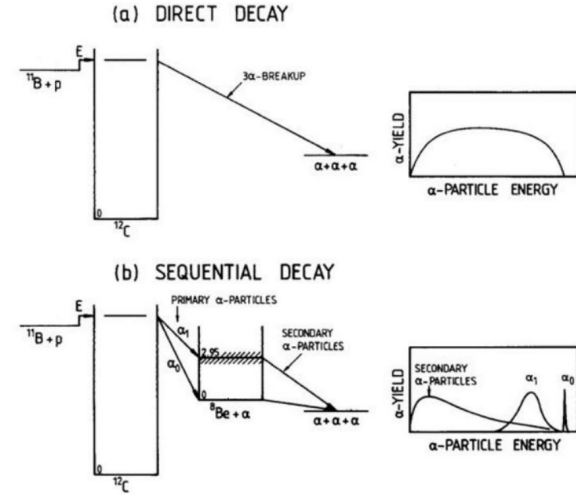
The mini - ELISSA chamber with the scaled down versions

Catalin Matei
Livio Lamia
Luca Guardo
Dimitar Balabanski
Teodora Petruse

.. and other members
of the collaboration
Involved in these
experiments.



Q-value = 8.59 MeV

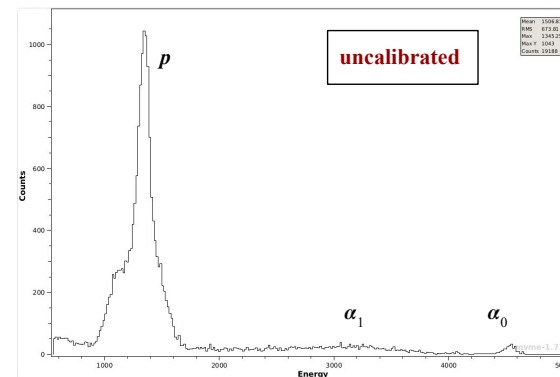
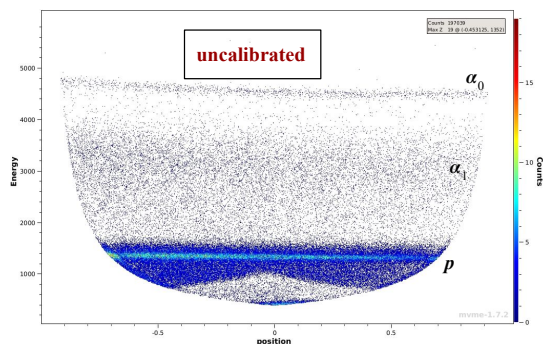
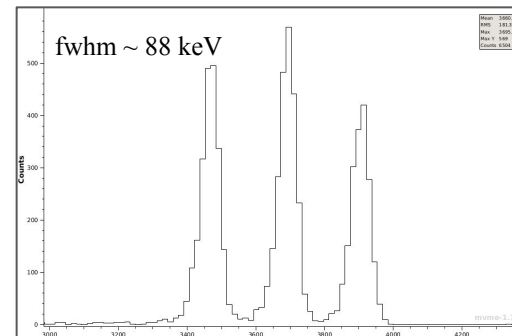
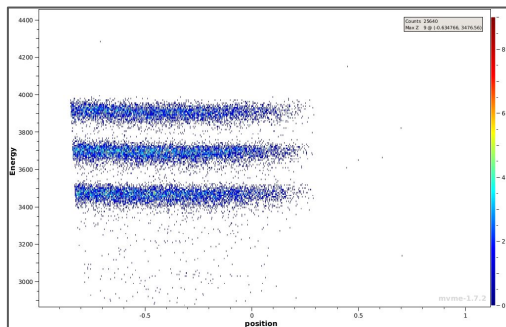


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Calibration

The X3 detector is a position sensitive detector (PSD). We calibrate these for energy using triple- α source for lower energy.

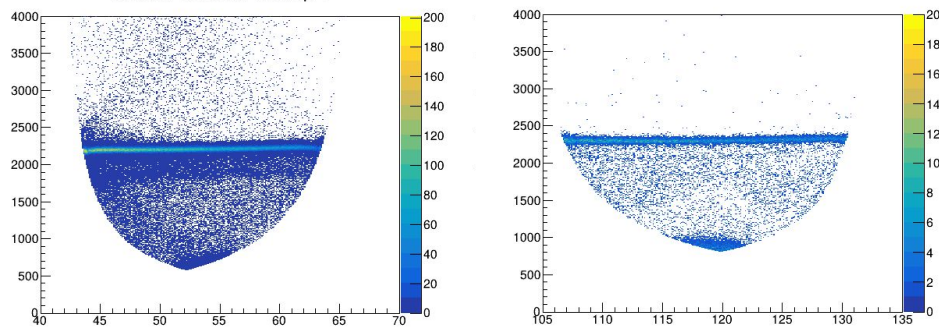
Position calibration is done using an equally spaced grid.



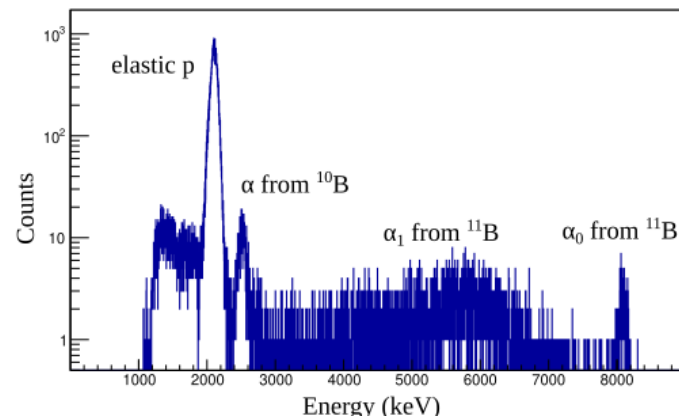
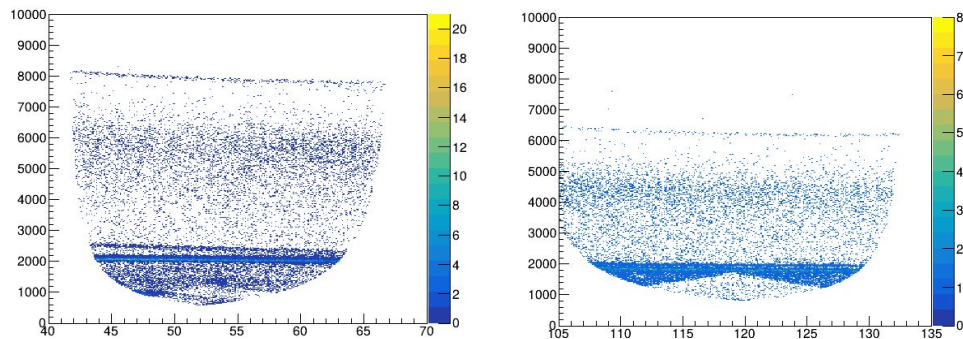
Raw spectra from p on ¹¹B @ E_p = 2500 keV

Energy-projection

E vs θ correlation for p+¹⁹⁷Au at 2.2 MeV

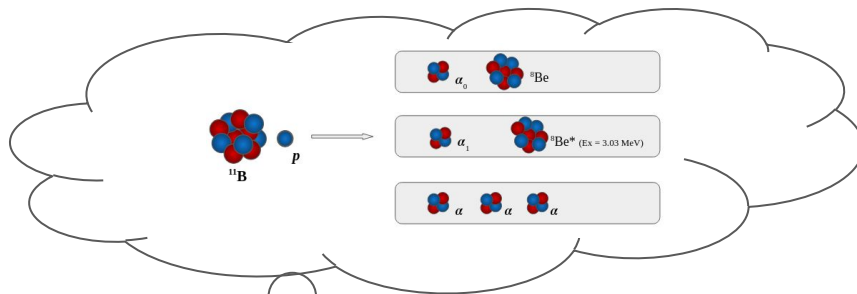


E vs θ correlation for p+¹¹B at 2.2 MeV



Energy spectra in one strip of a detector in the forward ring in the angular bin θ_{lab} : 45 - 46 deg. The peak at 2.2 MeV corresponds to the elastically scattered protons from ^{10,11}B. The next peak marked at 2500 keV is the α from reaction of p with ¹⁰B. We have also identified the α_0 peak around 8 MeV and the α_1 bump around 6 MeV from the transfer reaction.

- The ELISSA setup in its present configuration is very efficient in studying the charged particle reactions. The complete setup will be much more useful for detecting in **4 π configuration** and also higher efficiency in coincidence detection.
- The cross section for the energies ranging from **300 keV - 2600 keV** from a single experiment will provide a reliable set of parameter to analyze the reaction theoretically.
- The ¹¹B(p, α)⁸Be(2 α) experiment was performed in a **joint effort of ELI-NP and INFN-LNS**. The use of the **ELISSA** experimental setup highlighted once more its versatility for measuring low-energy nuclear reaction cross section.
- The preliminary data analysis is encouraging since it clearly shows the population of the ⁸Be ground state and its first excited state. Further improvements are still necessary and the work is in progress.



Thank you for your attention!