

New Measurement of the $^{27}\text{Al}(p,\alpha)^{24}\text{Mg}$ and $^{27}\text{Al}(p,\gamma)^{28}\text{Si}$ reaction cross section at LUNA

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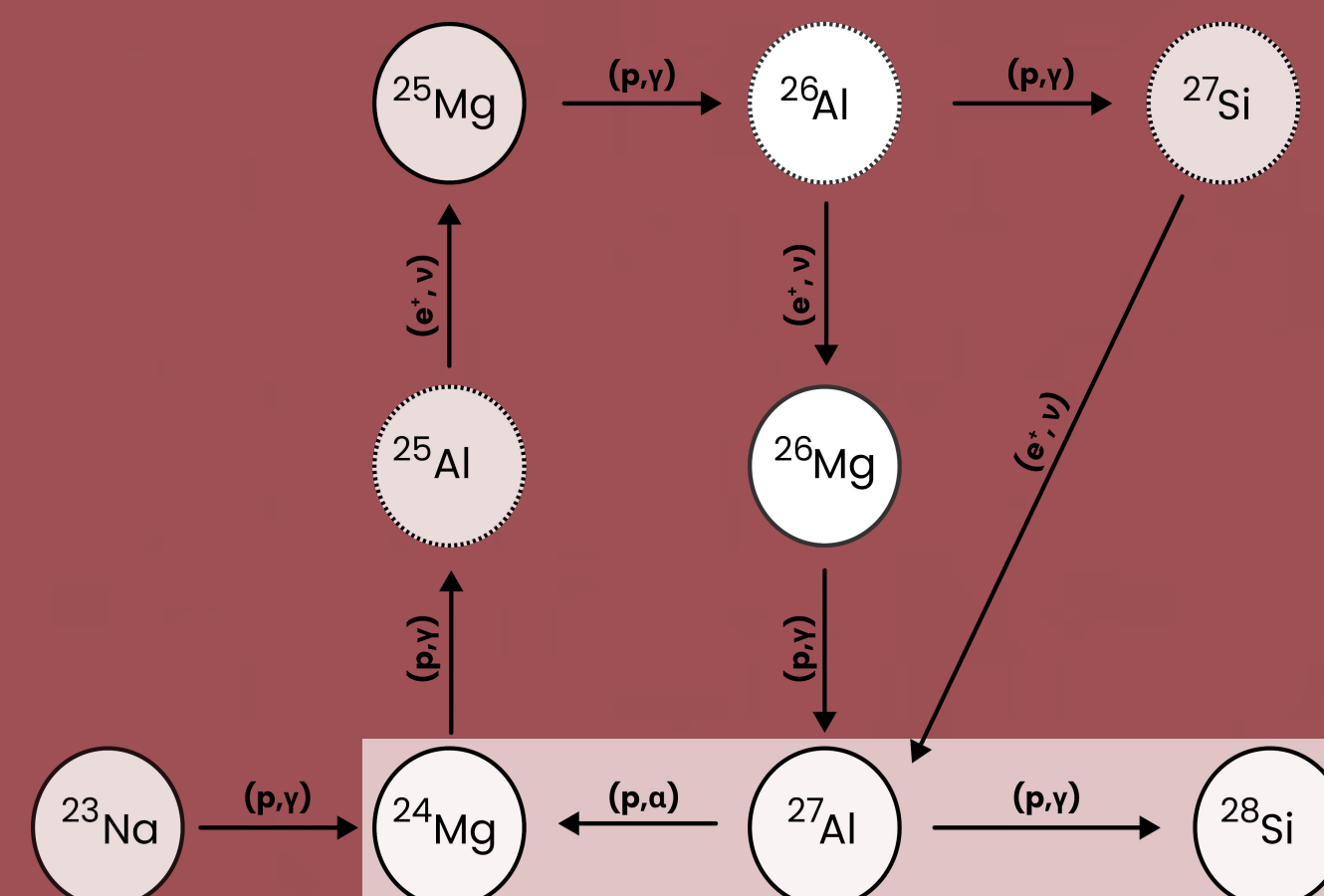
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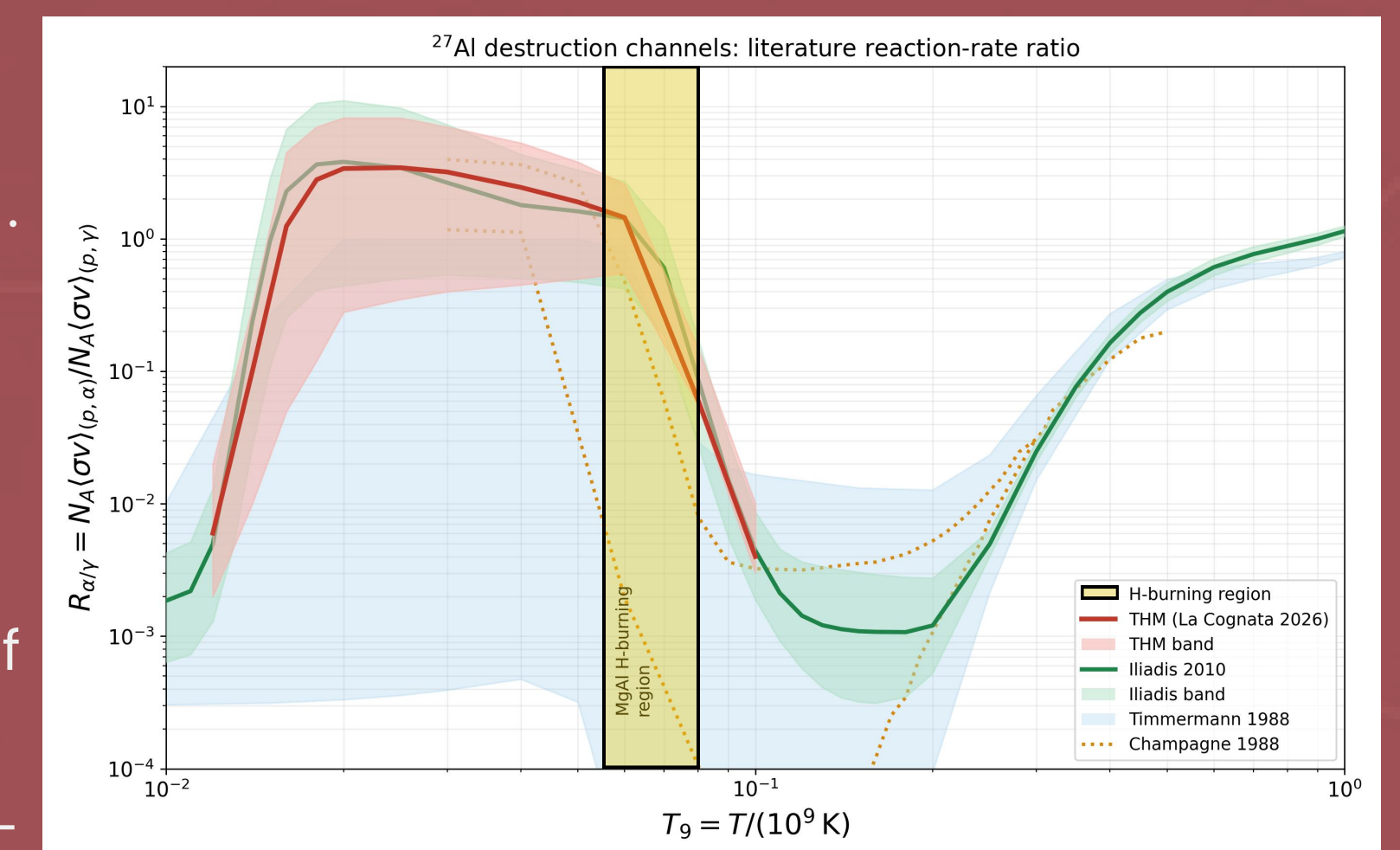
Astrophysical Motivation

^{27}Al is a branching point of the MgAl cycle. During stellar H-burning, ^{27}Al is destroyed through two competing channels: (p, α) and (p, γ) [1]. Their competition affects Mg and Al abundances, the interpretation of the Mg–Al anticorrelation observed in globular clusters [2] and ^{26}Al nucleosynthesis. The reaction rate ratio determines the closure of the Mg–Al cycle and the nucleosynthesis pathway to Si and P isotopes.



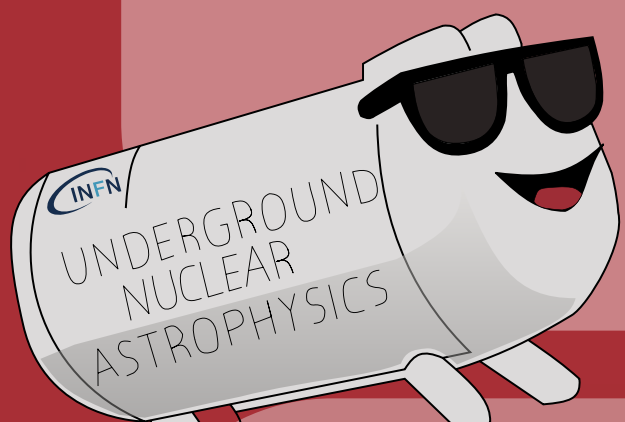
State of the Art

Low-energy resonances can dominate the stellar reaction rates, but their strengths are still poorly constrained [4]. Direct $^{27}\text{Al}(p,\gamma)^{28}\text{Si}$ resonance-strength measurements reach down to $E_p \sim 200$ keV [3], whereas previous direct $^{27}\text{Al}(p,\alpha)^{24}\text{Mg}$ searches mainly provided upper limits [5]. Direct measurements of the lowest accessible resonances are therefore crucial to reduce the reaction-rate uncertainties.

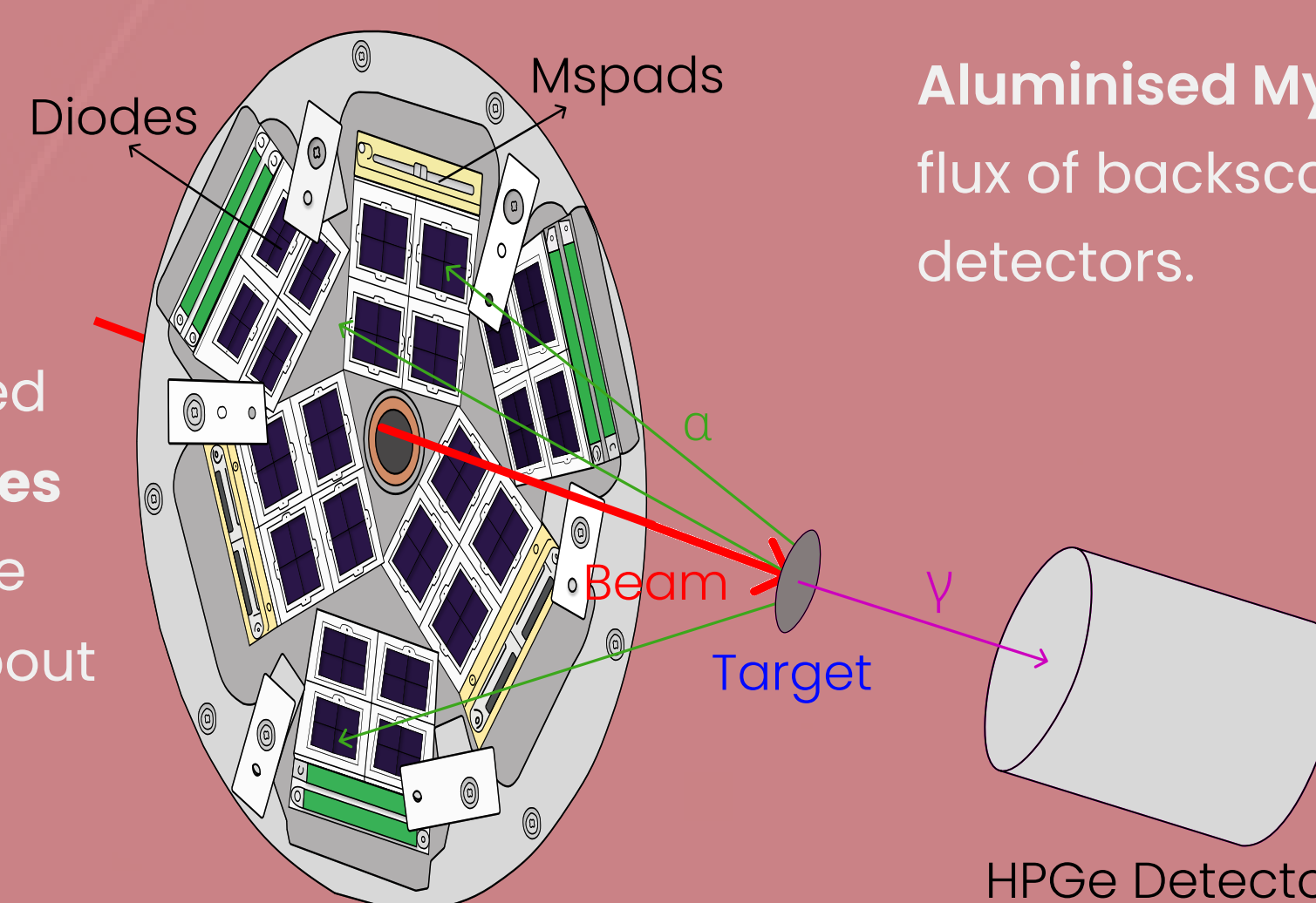


Experimental Setup

The experiment is performed at the **LUNA 400 kV** accelerator at INFN LNGS, where the underground location strongly suppresses the cosmic-ray background. A high-intensity proton beam (100–300 μA) impinges on a water-cooled ^{27}Al target.

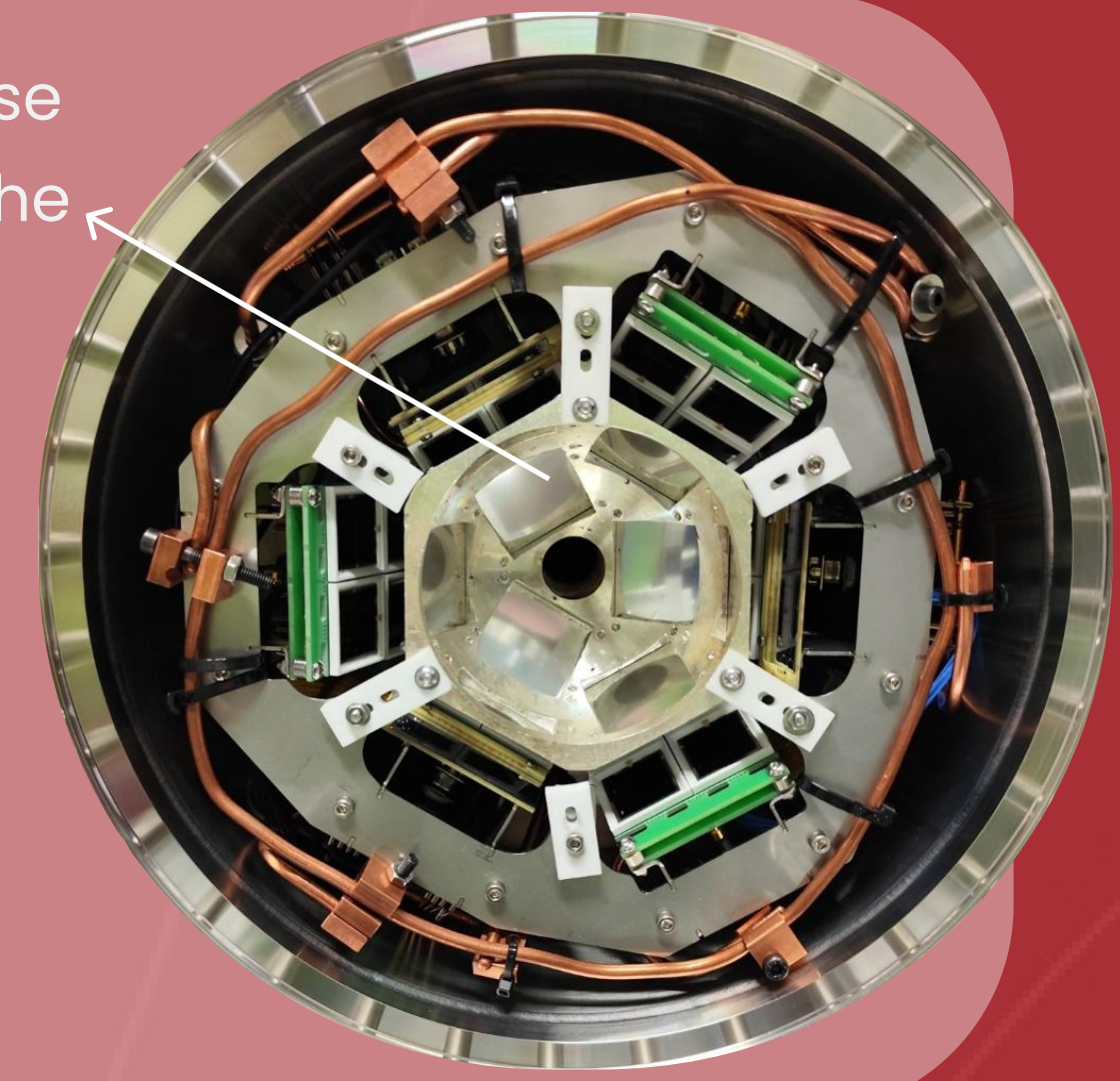


α -particles are detected at backward angles by a compact **Si array** composed of **24 Hamamatsu PIN diodes** and **3 MSPAD detectors**. The detectors are located at about 60 mm from the target.



Aluminised Mylar foils prevent the intense flux of backscattered protons reaching the detectors.

Gamma rays are detected by the **GeBochum HPGe detector**, positioned along the beam axis in close geometry.



Target Requirements

- **High stability** under intense proton irradiation;
- **Well-controlled thickness and reproducibility**, using targets with different thicknesses to assess systematic effects.
- High **chemical purity**, with particular attention to B contamination, which can generate a severe $^{10}\text{B}(p,\alpha)$ background in the charged-particle spectra.
- **Different target compositions and production techniques** to investigate potential systematic effects;

Backings & Cleaning

Ta PVD/ Ta backings

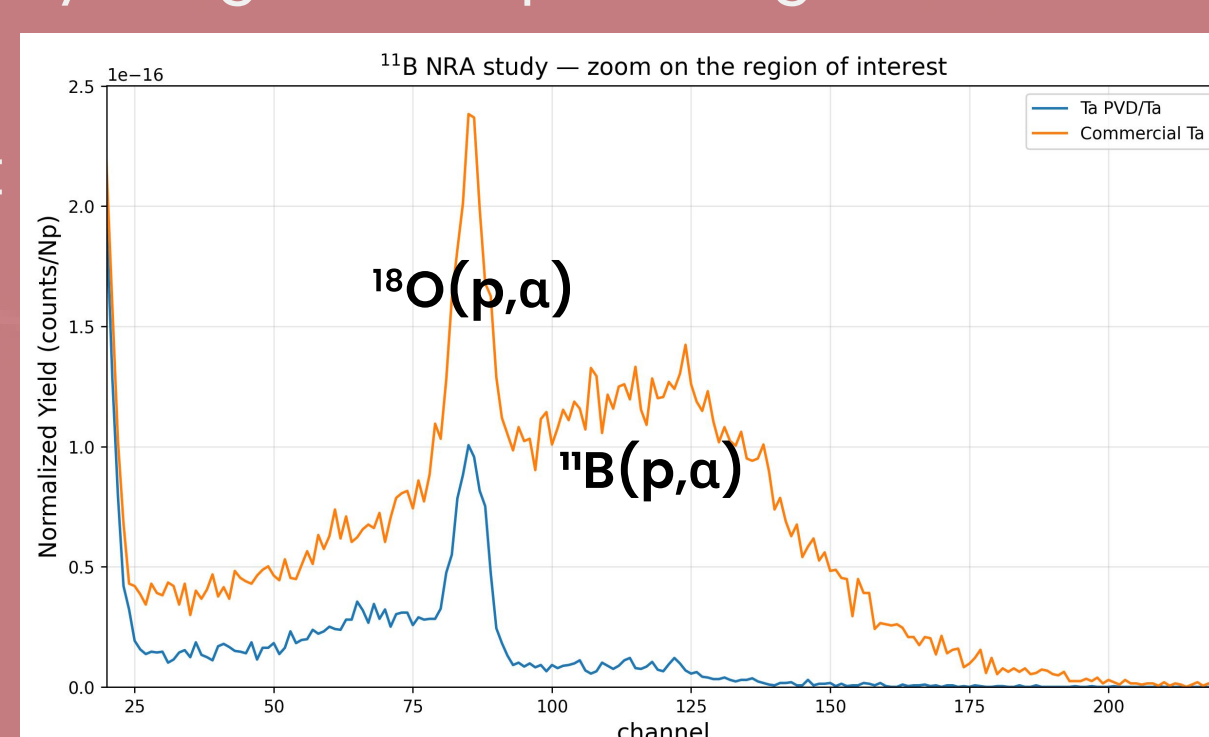
Commercial Ta substrates coated with a thin high-purity Ta layer by **Physical Vapor Deposition (PVD)** – INFN LNL Methodology

1. substrates cleaning [6]

- acetone + IPA $\rightarrow$ acid etching @ 80°C $\rightarrow$ DI-water rinse

2. PVD surface treatment: High-purity Ta deposited on the substrates by magnetron sputtering

$\rightarrow$ minimize B, F and light element contaminants.



Same cleaning procedure, different backing surface: Commercial Ta: 4.5×10^{13} ^{10}B atoms/cm² (~13 at. ppm) Ta PVD/Ta: 8.2×10^{12} ^{10}B atoms/cm² (~2 at. ppm)

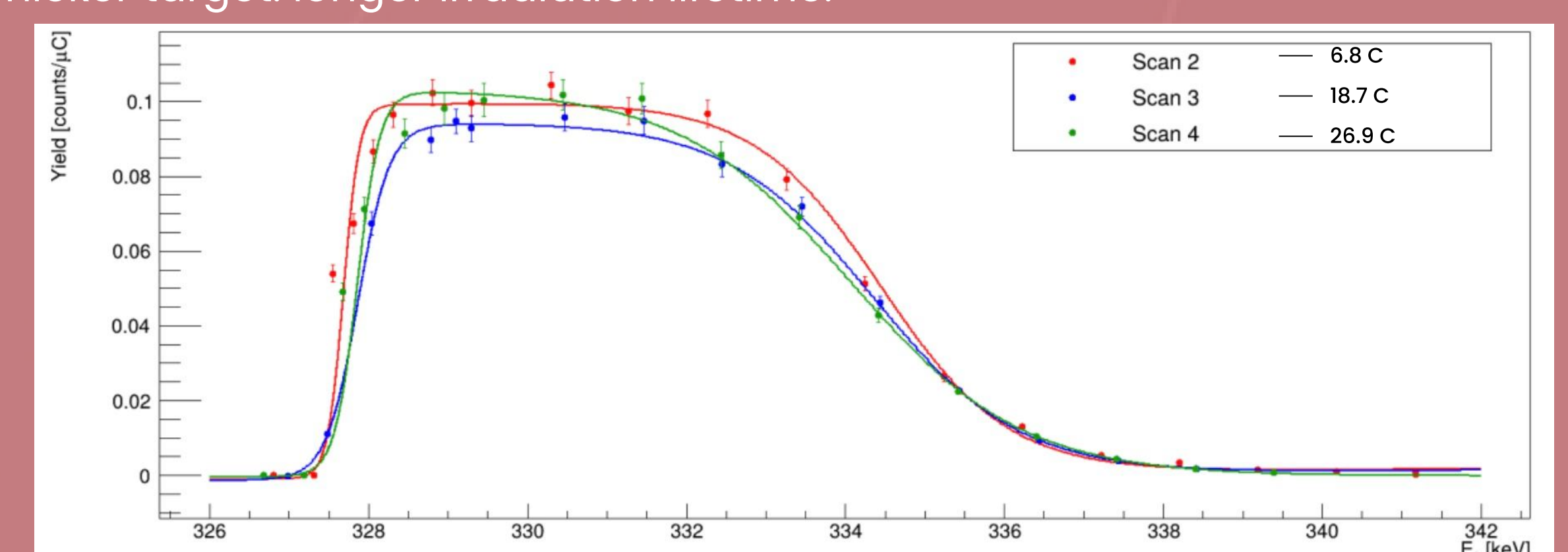
Target Candidates & First target Tests Results

- **Thermally evaporated Al/Ta targets – ATOMKI**
- **Commercial high-purity Al disks**
- **Magnetron-sputtered Al/Ta targets – INFN LNL**

Potentially improved control of thickness, uniformity and reproducibility.

Evaporated ^{27}Al targets produced at ATOMKI were tested through **resonance scans** after each long run in order to monitor changes in target thickness and resonance position.

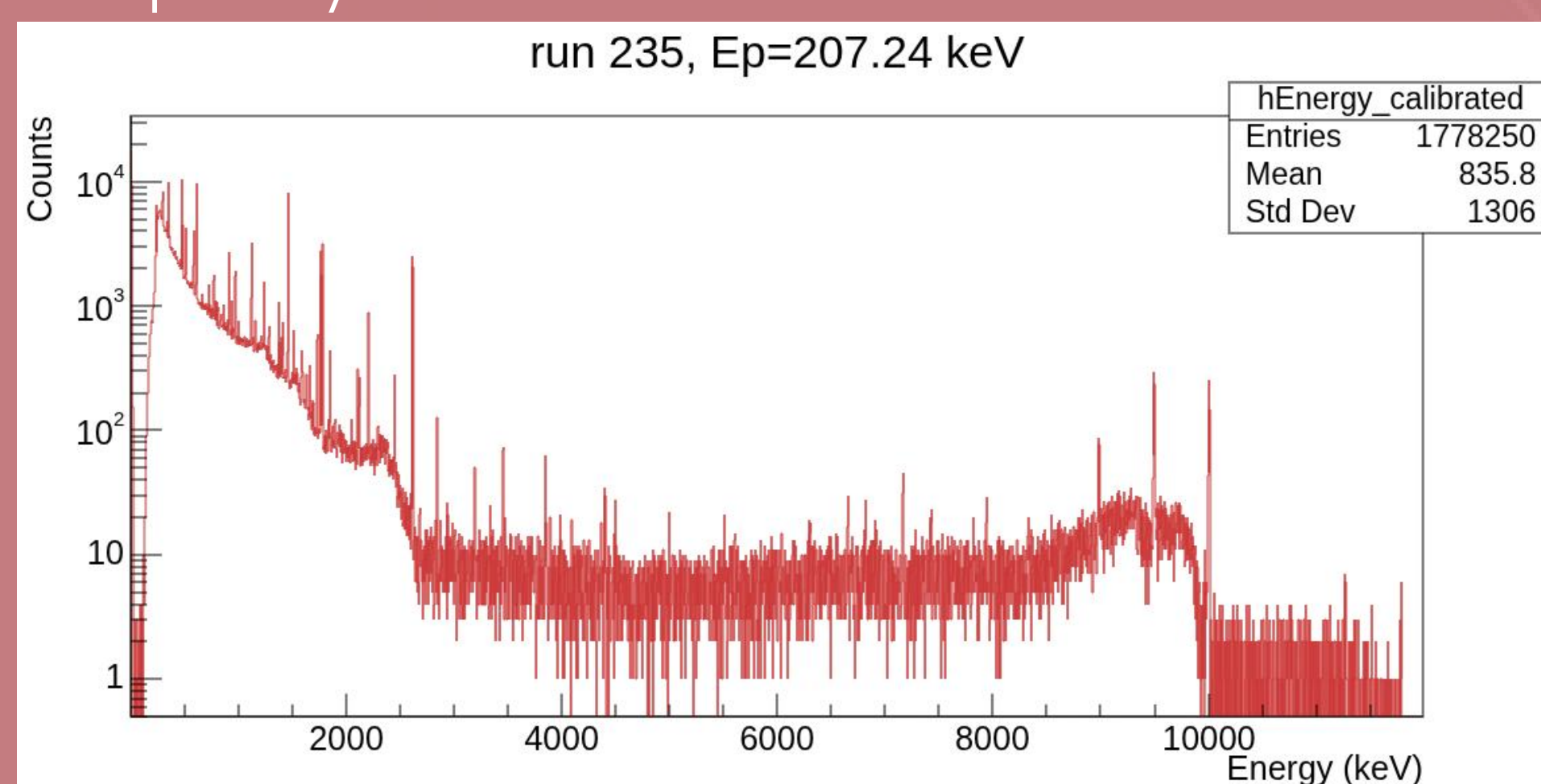
ATOMKI targets show stable yields under prolonged irradiation (up to 27 C). Thicker target: longer irradiation lifetime.



First Results – Sensitivity tests long runs were performed on top of all four resonances at $E_p \approx 207, 227, 296, 329$ keV.

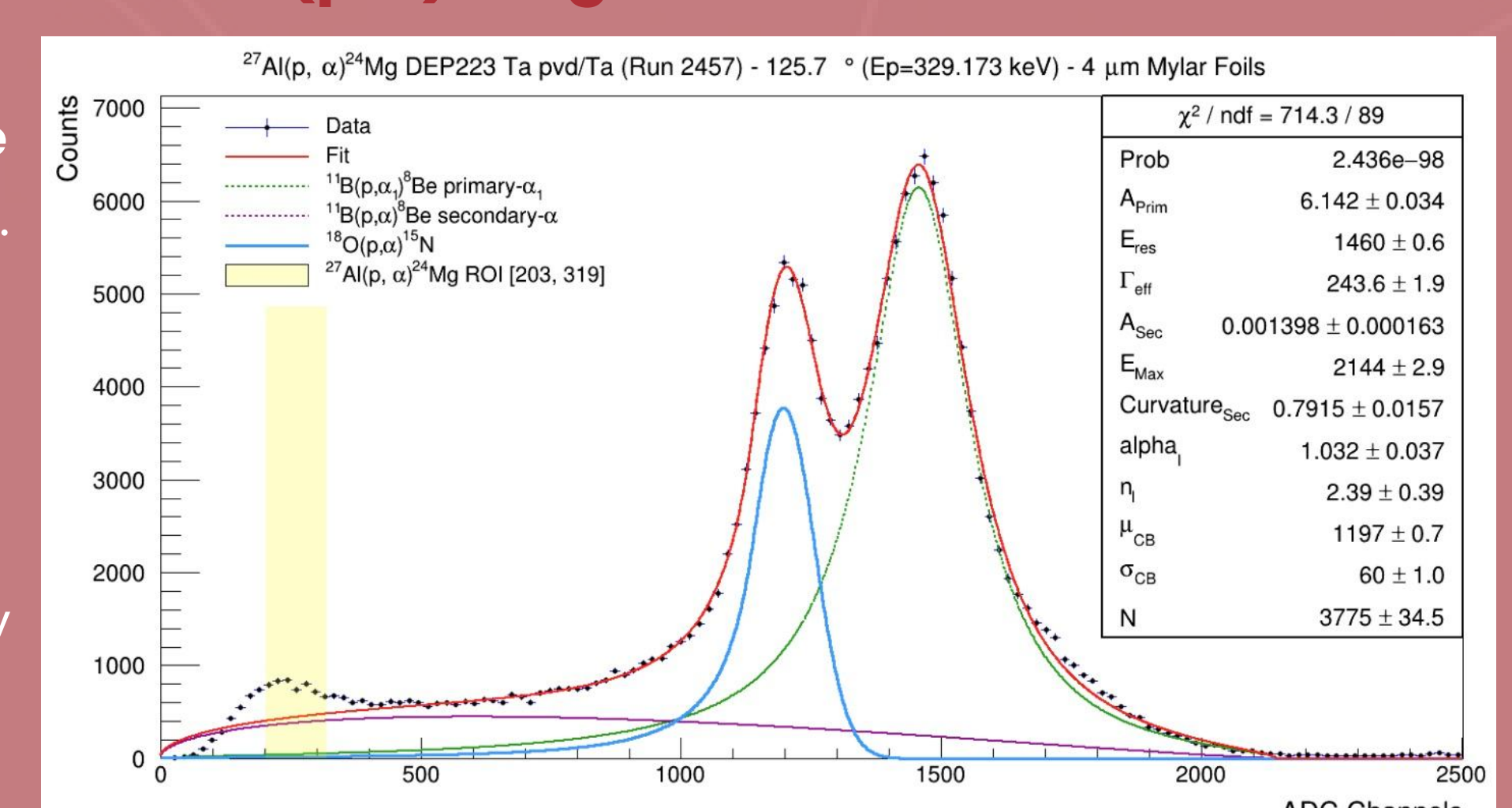
$^{27}\text{Al}(p,\gamma)^{28}\text{Si}$

Clear $^{27}\text{Al}(p,\gamma)^{28}\text{Si}$ signals were observed for all investigated resonances, including the expected primary transitions and additional transitions under investigation.



$^{27}\text{Al}(p,\alpha)^{24}\text{Mg}$

Preliminary long runs showed no clear evidence of the $^{27}\text{Al}(p,\alpha)^{24}\text{Mg}$ signal. The present sensitivity is limited by the low-energy background, in particular by ^{10}B contamination, detector noise and energy loss in the Mylar foils.



Improved electronics and thinner Mylar foils will be used to increase the (p, α) sensitivity

References:

- [1] A. E. Champagne, et al., Nuclear Physics A 487.2, 434–441 (1988). [4] M. La Cognata, EPJ Web Conf. 368, 00003 (2026). [6] Campostrini, et al., Eur. Phys. J. Spec. Top. 233, 2843–2857 (2024).
- [2] E. Carretta, et al., Astron. Astrophys., 505: 139–155 (2009). [5] R. Timmermann et al., Nucl. Phys. A 477, 105–119 (1988).