

THE IMPACT OF CARBON-OXYGEN SHELL MERGER ON INTERMEDIATE-MASS ELEMENTS IN MASSIVE STARS :

#79

INSIGHTS FROM IMPROVED CONVECTIVE BOUNDARY MIXING MODEL

WAN AISHAH HARUN¹, CLAUDIA LEDERER-WOODS¹, RAPHAEL HIRSCHI^{2,3}, MARCO PIGNATARI⁴, EMILY WHITEHEAD², UMBERTO BATTINO⁵, VISHNU VARMA²

¹ UNIVERSITY OF EDINBURGH, UK; ² KEELE UNIVERSITY, UK; ³ KAVLI IPMU, UNIVERSITY OF TOKYO, JAPAN; ⁴ KONKOLY OBSERVATORY, HUNGARY; ⁵ UNIVERSITY OF NAPLES FEDERICO II, ITALY.

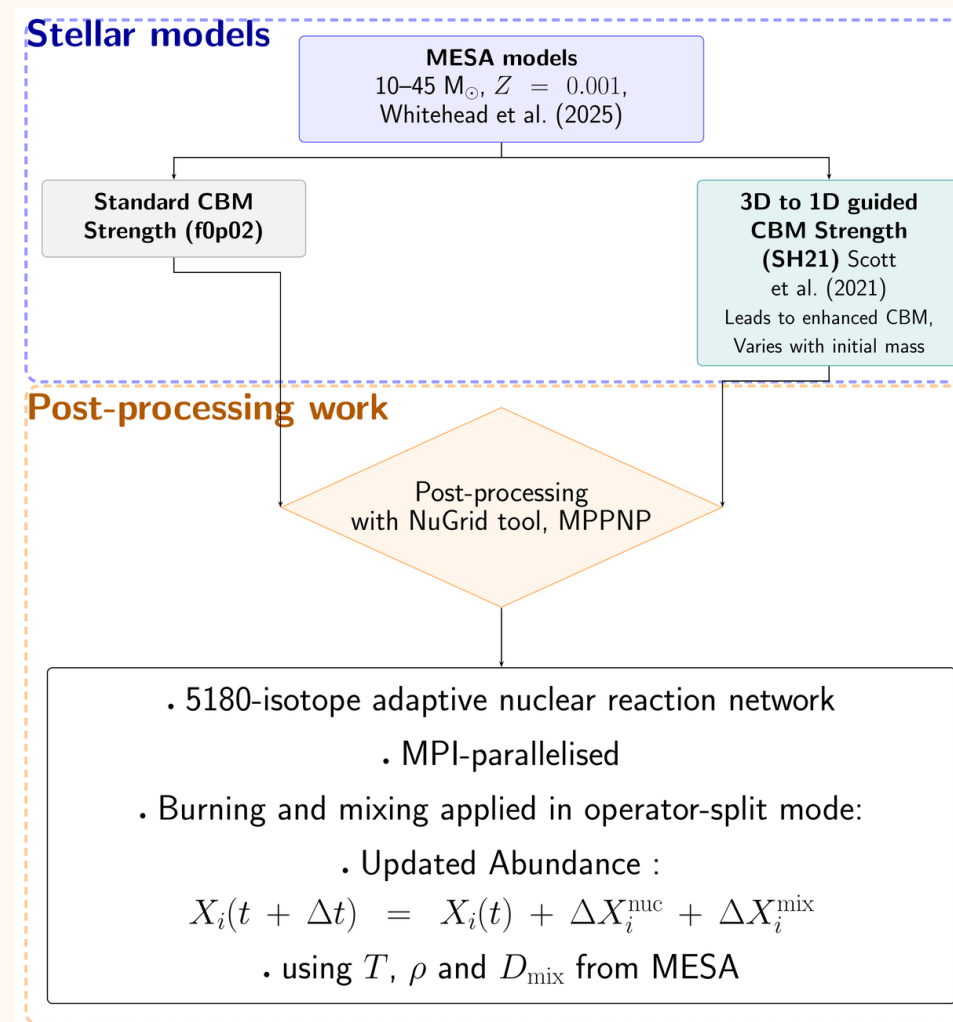
1 MOTIVATION

Standard massive-star models underproduce odd-Z elements (P, Cl, K, and Sc) at low metallicity compared with observations.

2 AIMS

Investigate the impact of the Convective Boundary Mixing (**CBM**) prescription motivated by 3D hydrodynamic simulations, on the role of **C-O shell merger** in the production of intermediate-mass elements.

3 METHODOLOGY



Frequency of C-O Shell Mergers :

C-O shell mergers occur in 85% of 3D-to-1D-guided CBM (**SH21**) and 95% of standard CBM models (**f0p02**).

4 KEY FINDINGS

Odd-Z Production:

- Enhanced P, K, Cl, and Sc production in **3D-to-1D-guided CBM (SH21) with C-O shell mergers**.
- Up to 10× higher than in standard CBM (f0p02) with C-O shell mergers.

Role of Peak Temperature:

Peak temperature during shell mergers influences the production of elements around neon and magnesium.

